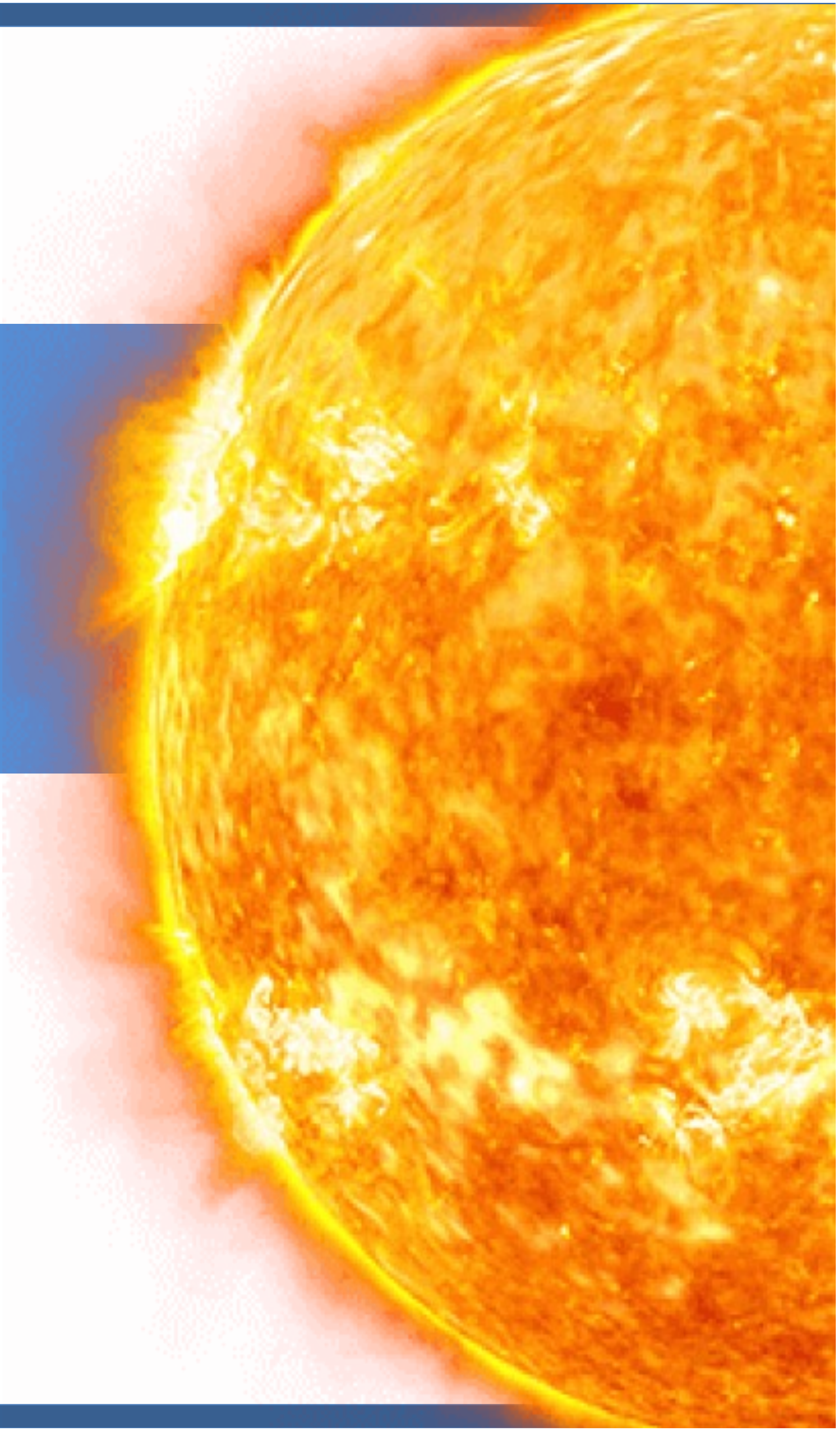
X-ray DEtector Technologies for Physics (XDEP)

Status of IAXO & BabyIAXO experiment

Jaime Ruz on behalf of the IAXO Collaboration

February 5-6, 2024

SOLEIL, France



CAPA Centro de Astropartículas y Física de Altas Energías
Universidad Zaragoza

Outline

1. Standard Model, Strong CP and Dark Matter

2. The Axion

3. Detection of Axions. Solar Axion Searches

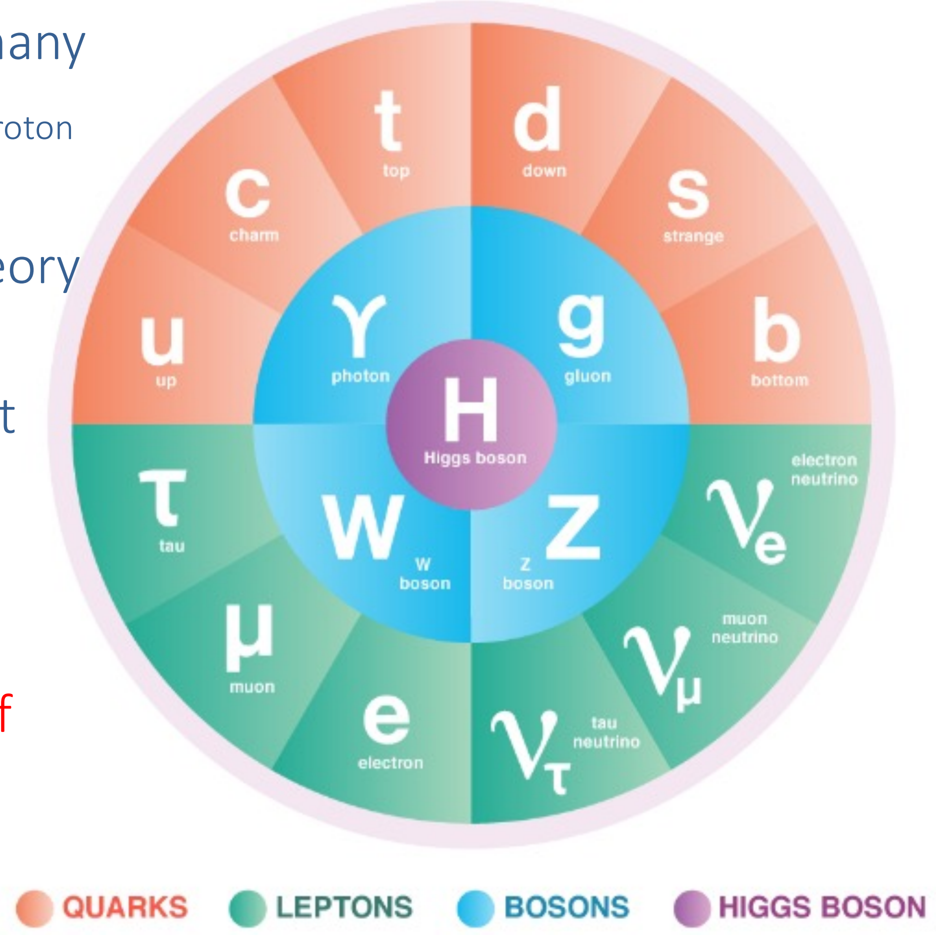
4. IAXO and BabyIAXO

5. Conclusions

Standard Model, strong CP

STANDARD MODEL (SM) OF PARTICLE PHYSICS

- ✓ Extremely successful theory describing many observations up to energies of ~1000 m_{proton}
- ✓ Merely an effective theory that could be considered the low energy limit of a Theory of Everything
- ✓ Expect observation of new phenomena at higher energies (e.g. LHC at CERN)
- ✓ SM cannot explain:
  - What is the nature of dark matter?
  - Why is the electric dipole moment of the neutron so small?



Neutron electric dipole moment 2nd Challenge

Why is the nEDM so small?

- ✓ QCD Lagragian contains a CP violating term (with q-parameter of QCD vacuum)

$$\mathcal{L}_{\not{CP}} = \overline{\theta}\frac{\alpha_s}{8\pi}G^a_{\mu\nu}\widetilde{G}^{\mu\nu}_a$$

- ✓ Observational Consequences: Prediction of electric dipole moments (EDM) to hadrons, most importantly, to neutrons

$$d_n \sim 10^{-16}\ \bar{\theta}\ e\,\mathrm{cm}$$

Crewther,Di Vecchia,Veneziano,Witten 1979;...; Pospelov,Ritz 2000

- ✓ Latest measurements of the nEDM

$$|d_n| < 1.8 \times 10^{-26}\, e\,\mathrm{cm}$$

Abel et al. 2020

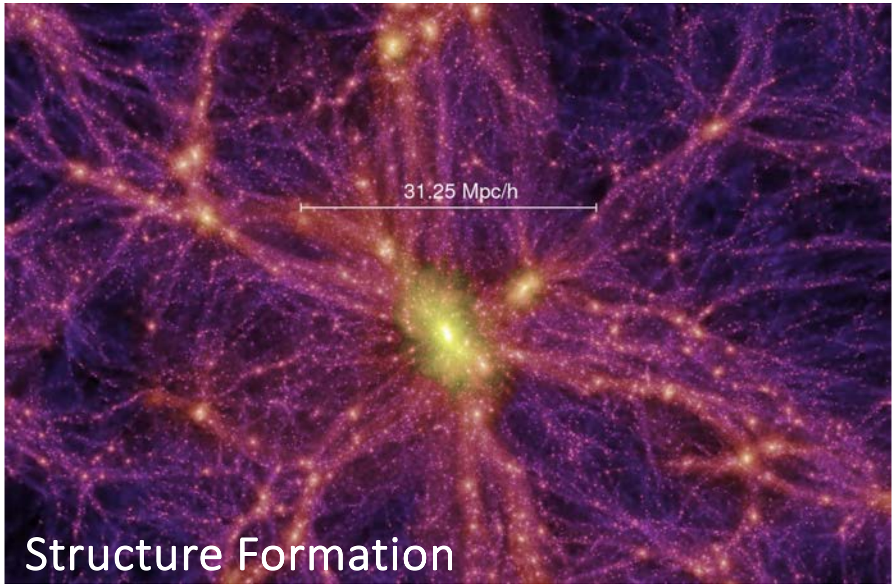
- ✓ Therefore expect $|\overline{\theta}| \lesssim 10^{-10}$

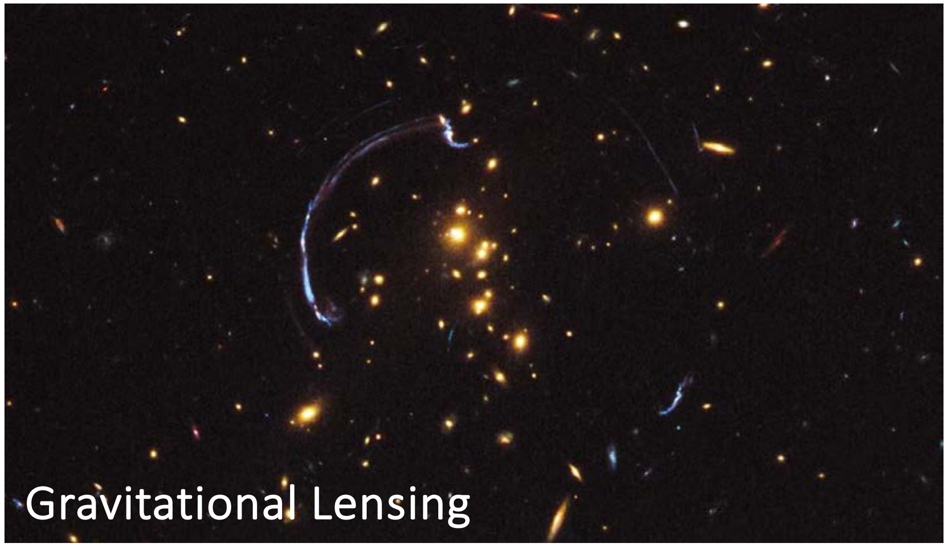
STRONG CP PROBLEM or WHY IS THETA SO SMALL?

Dark Matter

EVIDENCE FOR DARK MATTER

- ✓ Galaxy rotation curves
- ✓ Cosmic Microwave Background (CMB)
- ✓ Structure formation
- ✓ Gravitational lensing
- ✓ Bullet Cluster





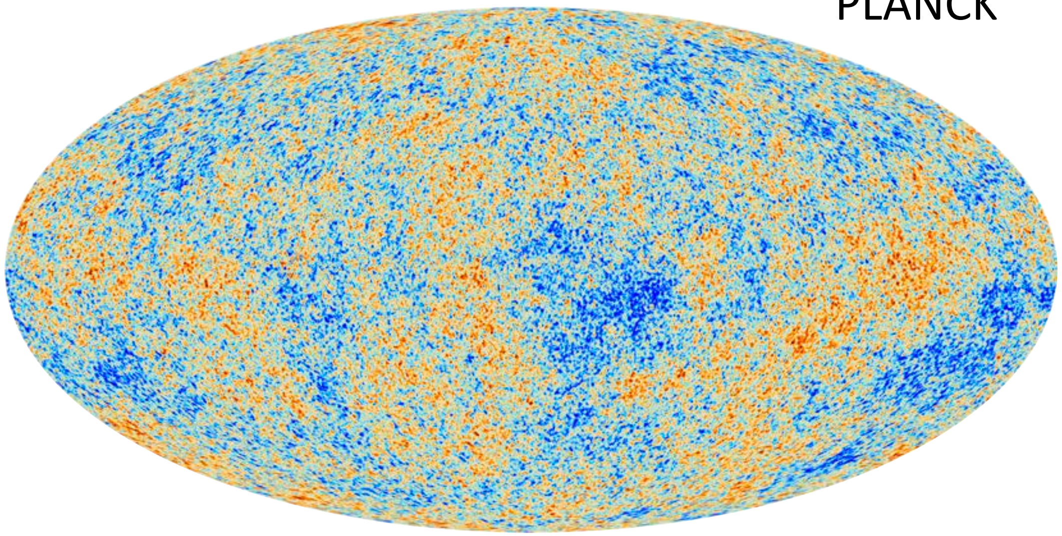
DARK MATTER PROBLEM: WE KNOW IT EXISTS BUT WHAT IS ITS NATURE?

Dark Matter

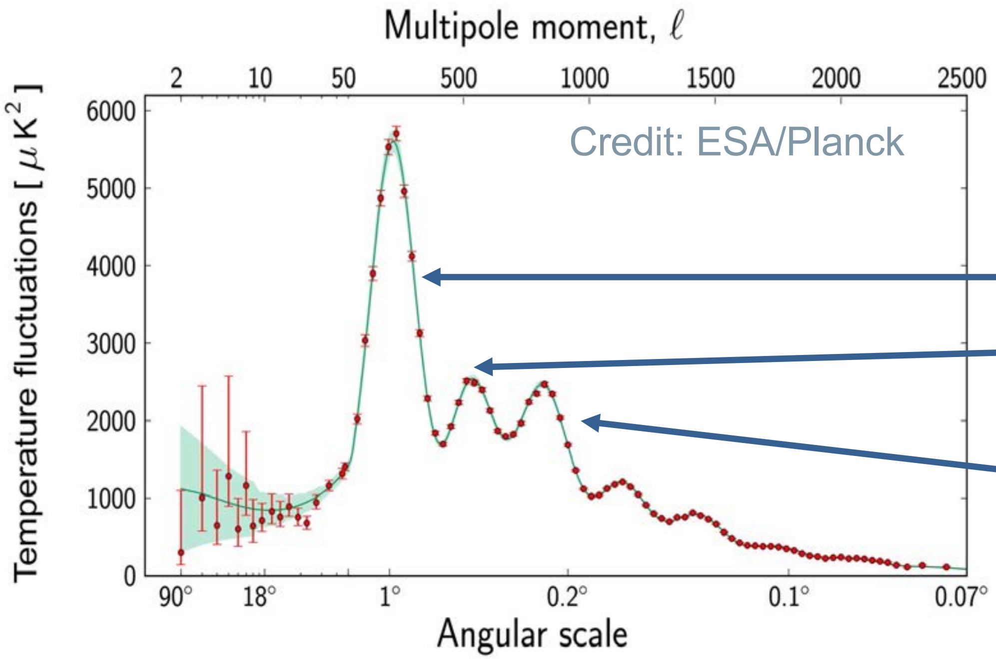
EVIDENCE FOR DARK MATTER

- ✓ Galaxy rotation curves
- ✓ Cosmic Microwave Background (CMB)

PLANCK



PLANCK power spectrum of the CMB radiation temperature anisotropy



Location and height of peaks determines cosmological parameters

Flat universe

Baryonic matter ~5% of the total mass/energy of the universe

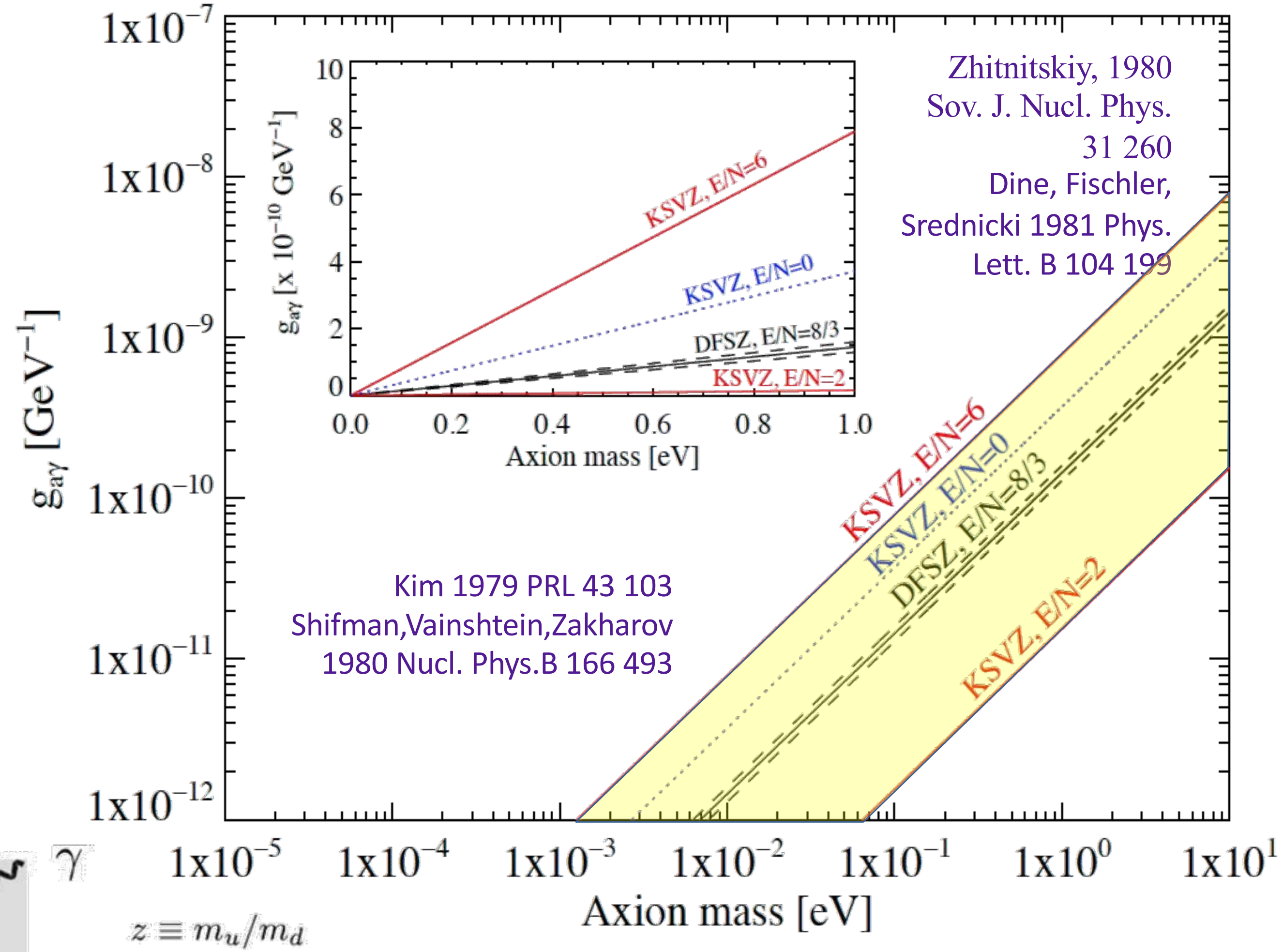
Amount of **dark matter** ~27%

The Axion

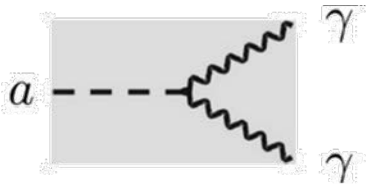

Standard Model
AXION
HYPOTHETICAL ELEMENTARY PARTICLE
PROMISES TO
"CLEAN UP"
YOUR STRONG
CP PROBLEM!

The Axion

$$\mathcal{L} \supset \frac{g_{a\gamma\gamma}}{4}\, a\, F_{\mu\nu}\tilde{F}^{\mu\nu} \equiv g_{a\gamma\gamma}\, a\, \mathbf{E}\cdot\mathbf{B}$$



$$g_{a\gamma} \simeq \frac{\alpha}{2\pi f_\pi}\frac{m_a}{m_\pi}\frac{1+z}{\sqrt{z}}\left(\frac{E_Q}{N_Q} - \frac{2}{3}\frac{4+z}{1+z}\right)$$

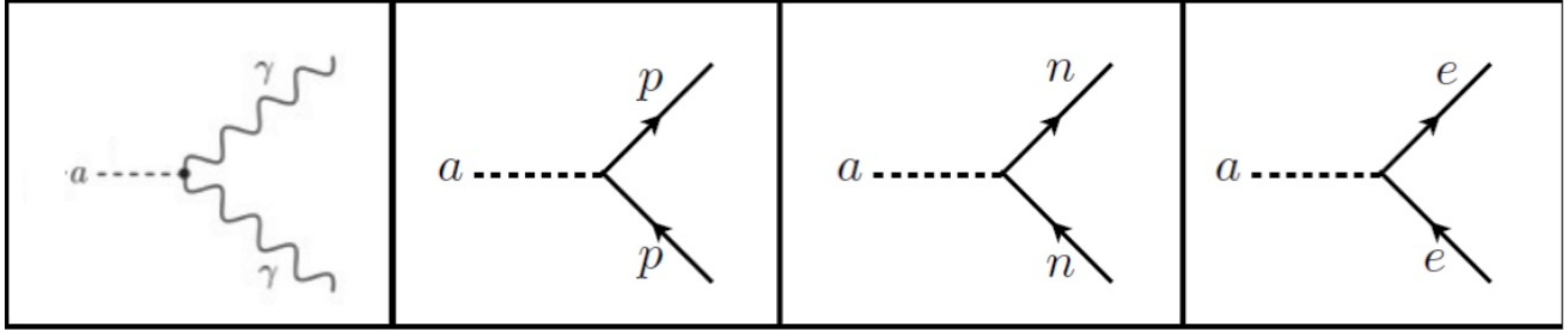


$z \equiv m_u/m_d$

AXION COUPLINGS

- ✓ Axions interact with photons (generic) and with regular matter (model dependent)

2 photon	proton	neutron	electron
$\frac{\alpha C_{a\gamma}}{2\pi}\frac{a}{f_a}\frac{F_{\mu\nu}\widetilde{F}^{\mu\nu}}{4}$ +	$C_{ap} m_p \frac{a}{f_a}[i\bar{p}\gamma_5 p]$ +	$C_{an} m_n \frac{a}{f_a}[i\bar{n}\gamma_5 n]$ +	$C_{ae} m_e \frac{a}{f_a}[i\bar{e}\gamma_5 e]$ +



$$g_{a\gamma} = \frac{C_{a\gamma}\alpha}{2\pi f_a} \qquad g_{ap} = C_{ap}\frac{m_p}{f_a} \qquad g_{ap} = C_{an}\frac{m_n}{f_a} \qquad g_{ap} = C_{ae}\frac{m_e}{f_a}$$

- ✓ Due to its properties axions are favored dark matter candidates (next to WIMPs)

Detection of Axions

Detection techniques

Source		Experiments	Model & cosmology dependency
Relic axions		Haloscopes	High (assumes axions are all of the DM)
Lab axions		Light-Shining-Through-Wall Experiments	Very low
Solar axions		Helioscopes	Low

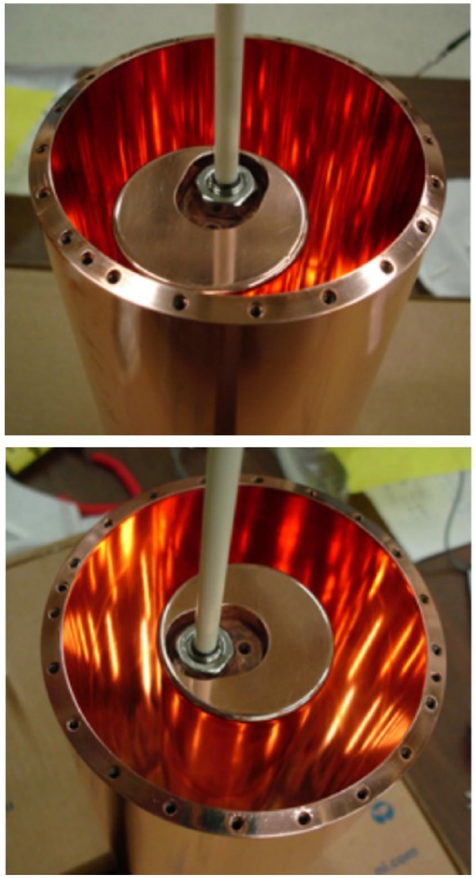
Large complementarity between different experimental approaches!

Detection of Axions

EXPERIMENTS RELYING ON AXIONS BEING DARK MATTER

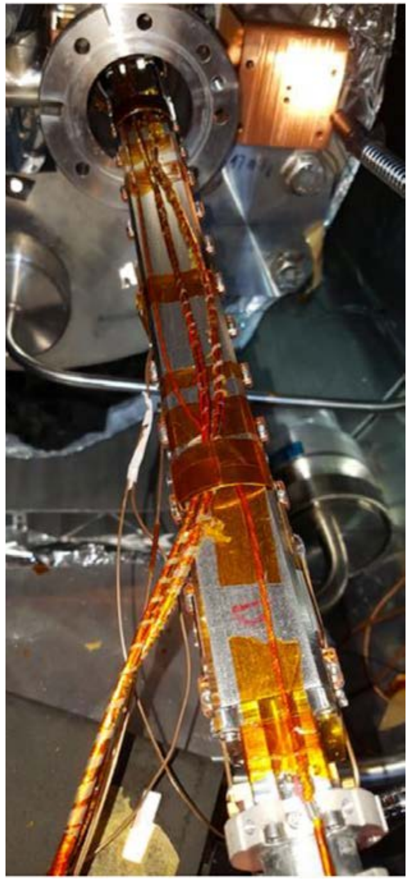
ADMX

HAYSTAC



RADES

CAST-CAPP



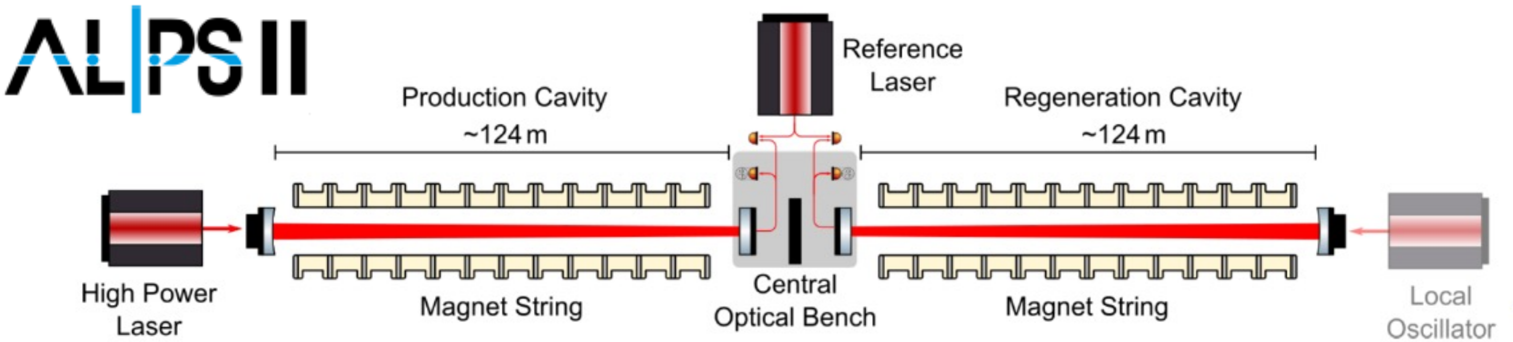
M. Maroudas et al 2022 Nature Com. 13 1 1-9
Alvarez et al 2021 JHEP 2021 75

Detection of Axions — Shining-light-through walls

EXPERIMENTS NOT RELYING ON AXIONS BEING DARK MATTER

- LIGHT-SHINING-THROUGH-WALL EXPERIMENTS: pure laboratory searches
  - ✓ ALPS

    Most basic layout of a LSTW experiment
  - ✓ ALPS-II
    - 12 + 12 straightened HERA magnets
    - Optical cavities both at production and regeneration sites
    - Sensitivity 3000×ALPS



Ringwald 2003 Phys. Lett. B 569 51

Detection of Axions

Helioscopes

EXPERIMENTS NOT RELYING ON AXIONS BEING DARK MATTER

- AXION HELIOSCOPES: laboratory axion searches looking for solar axions

CAST Collaboration 2017 Nature Phys. 13 584-590
Arik et al 2015 PRD 92 021101
Arik et al 2014 PRL 112 091302
Barth et al 2013 JCAP 1305 010
Arik et al2011 PRL107 261302
Zioutas et al 2009 JCAP 0902 008
Zioutas et al 2007 JCAP 0704 010

CERN AXION SOLAR TELESCOPE (CAST)

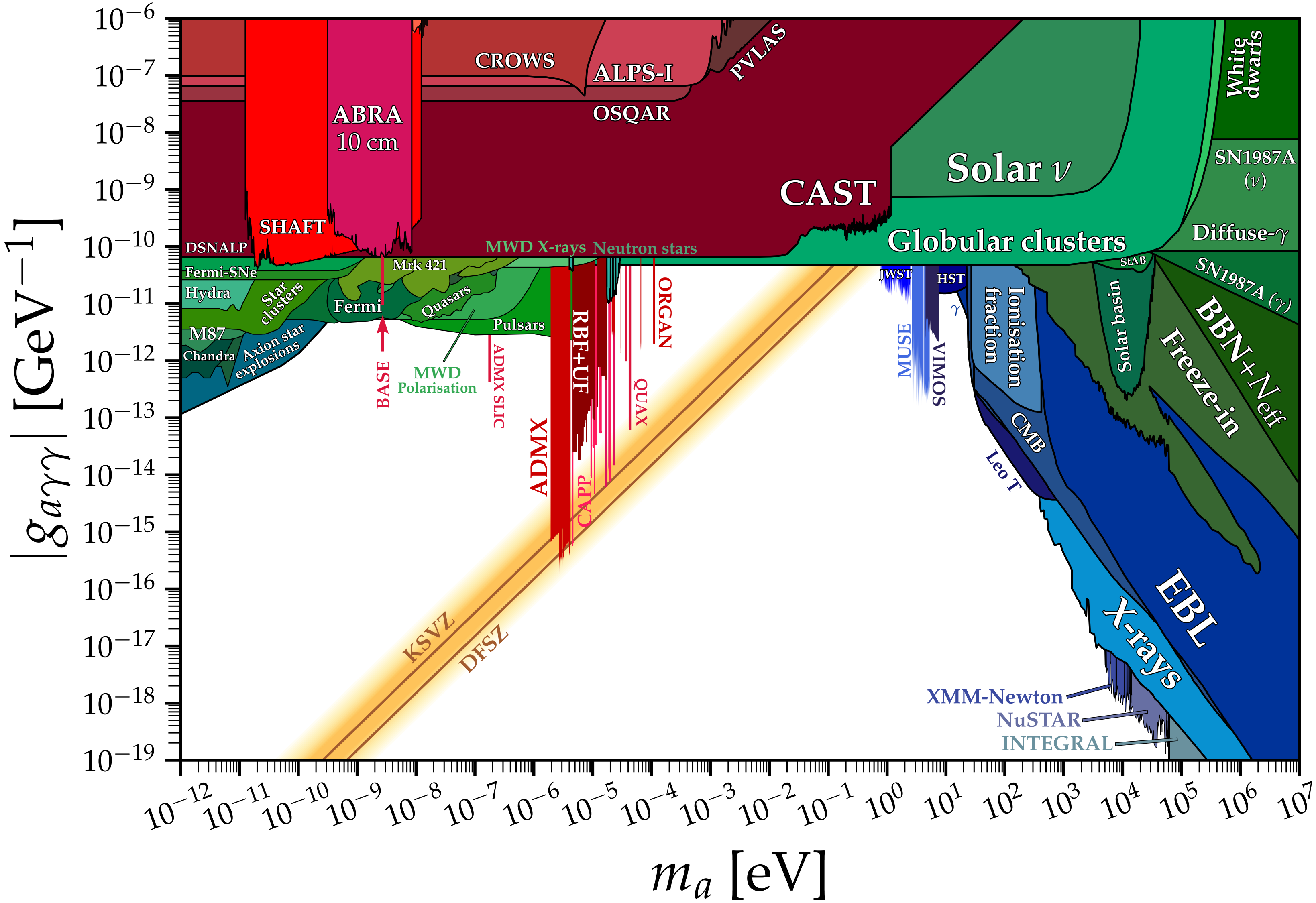
- Most powerful axion helioscope to date
- Superconducting prototype LHC dipole magnet
- X-ray focusing devices and ultralow-background detectors
- Use of buffer gas to extend sensitivity to higher masses (axion band)

STATE-OFF-THE-ART … SO FAR …

Detection of Axions

State-of-the-art

Adapted from https://github.com/cajohare/AxionLimits

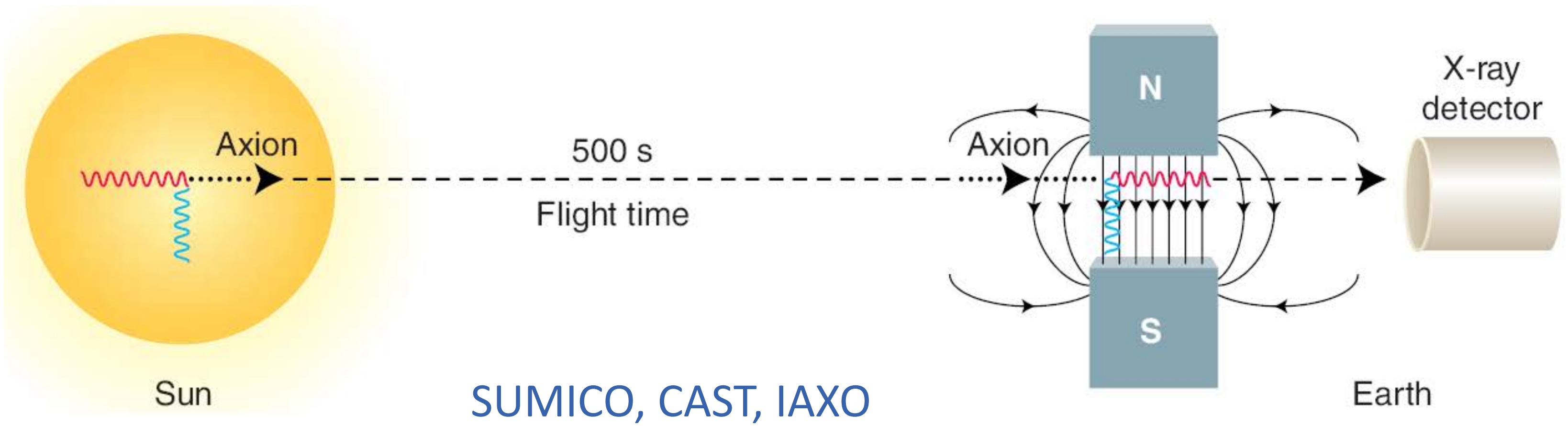


Solar Axion Searches

EXPERIMENTS NOT RELYING ON AXIONS BEING DARK MATTER

- AXION HELIOSCOPES: laboratory axion searches looking for solar axions

P. Sikivie 1983 PRL 51 1415



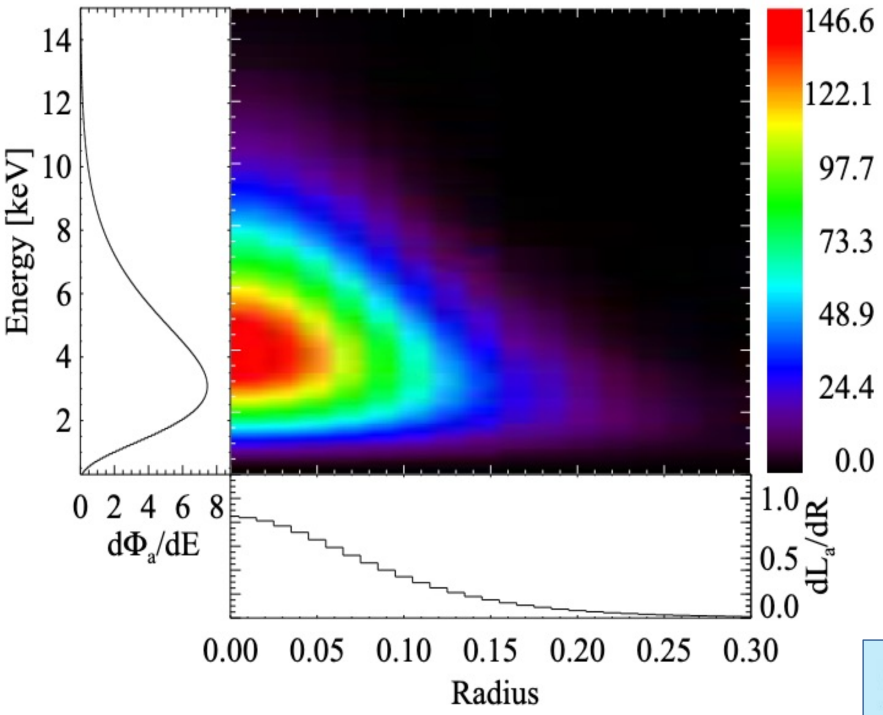
SUMICO, CAST, IAXO

Concept:

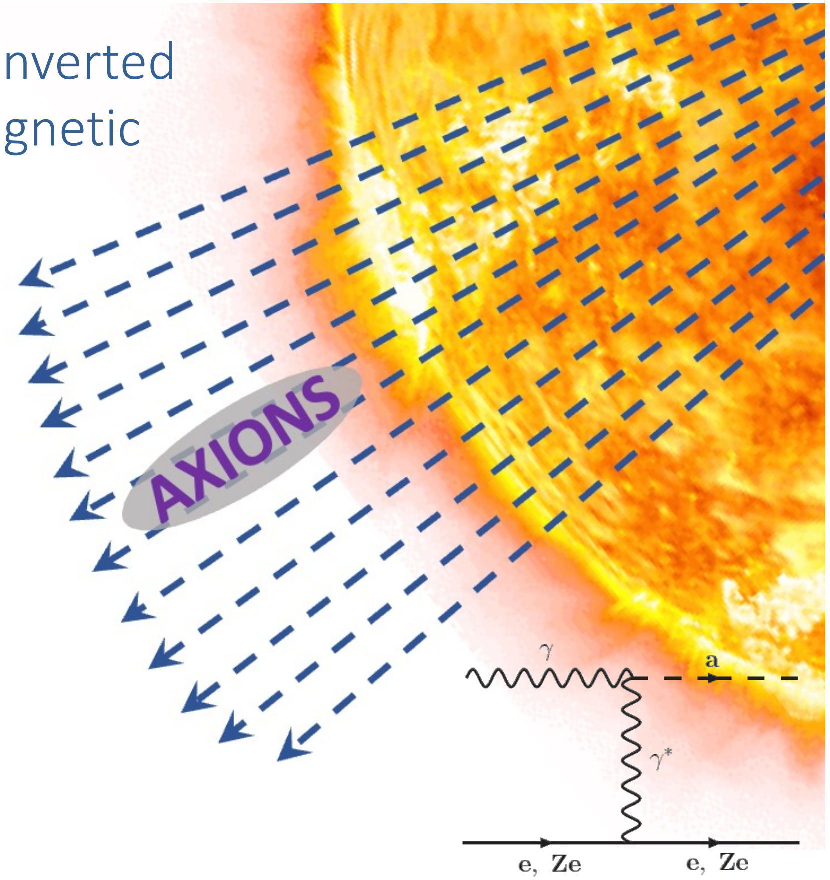
- Axions produced in strong electromagnetic fields of the core of the Sun
- Solar axion conversion into x-ray (keV) photons in transverse laboratory B-field

Solar Axion Searches

- Blackbody photons (keV) in solar core can be converted into axions in the presence of strong electro magnetic fields in the plasma → Primakoff Effect.



Hadronic axions (if the axion couples predominantly to photons (g_{ag}))



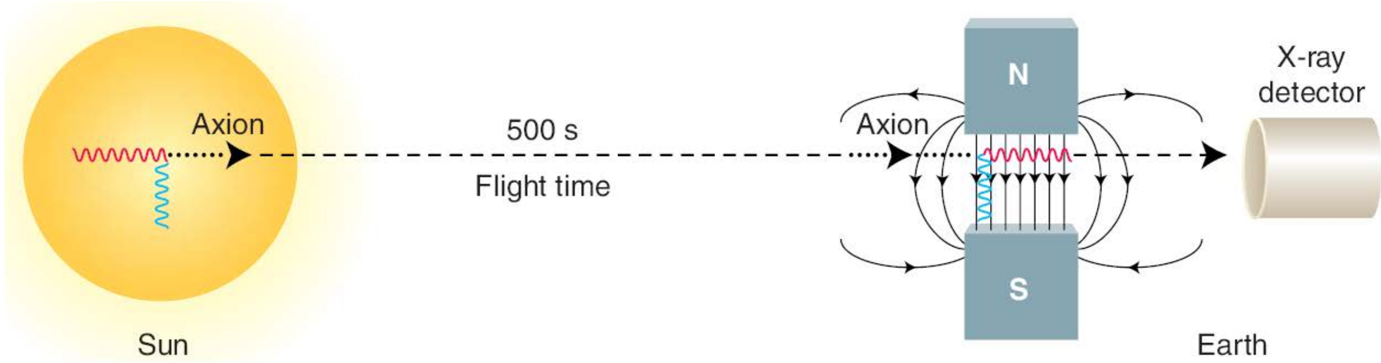
$$\frac{d\Phi_a}{dE} = 6.02 \times 10^{10} \left(\frac{g_{a\gamma}}{10^{-10}\mathrm{GeV}^{-1}} \right)^2 E^{2.481} e^{-E/1.205} \frac{1}{\mathrm{cm}^2\ \mathrm{s}\ \mathrm{keV}}$$

Solar Axion Searches — Detection Helioscope

- First axion helioscope proposed by P. Sikivie

P. Sikivie 1983 PRL 51 1415

Reconversions of axions into x-ray photons possible in strong laboratory magnetic field



$$P_{a\to\gamma} = \left(\frac{BLg_{a\gamma\gamma}}{2}\right)^2 \quad \text{for} \quad \frac{qL}{2} < \pi \quad \text{with} \quad q = \frac{m_a^2}{2E_a}$$

VACUUM

- Idea refined by K. van Bibber et al.

Van Bibber et al 1989 Phys. Rev. D 39 2089

Buffer gas to restore coherence over long magnetic field and access higher axion masses

$$P_{a\to\gamma} = \left(\frac{Bg_{a\gamma\gamma}}{2}\right)^2 \frac{1}{q^2+\Gamma^2/4}\left[1+\mathrm{e}^{-\Gamma L} - 2\mathrm{e}^{-\Gamma L/2}\cos(qL)\right] \quad \text{with} \quad q = \left|\frac{m_\gamma^2 - m_a^2}{2E_a}\right|$$

GAS

Solar Axion Searches

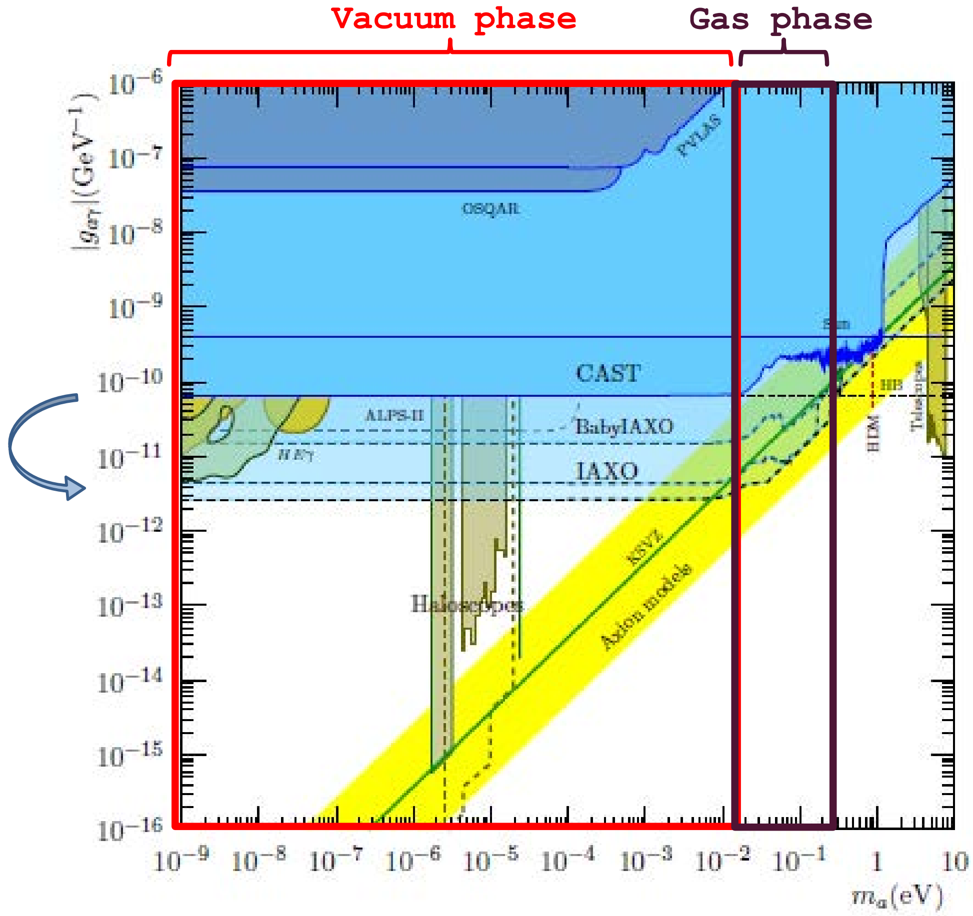
Detection Helioscope

- Vacuum Phase:
  - Coherence condition valid for $m_a \lesssim 0.02$ eV
- Gas Phase:
  - Extends coherence condition valid from $0.02 \text{ eV} \lesssim m_a \lesssim 0.26$ eV

$$m_\gamma = 4.498716\sqrt{\frac{P_{He}[\text{atm}]}{T_{He}[\text{K}]}}\,\text{eV}.$$

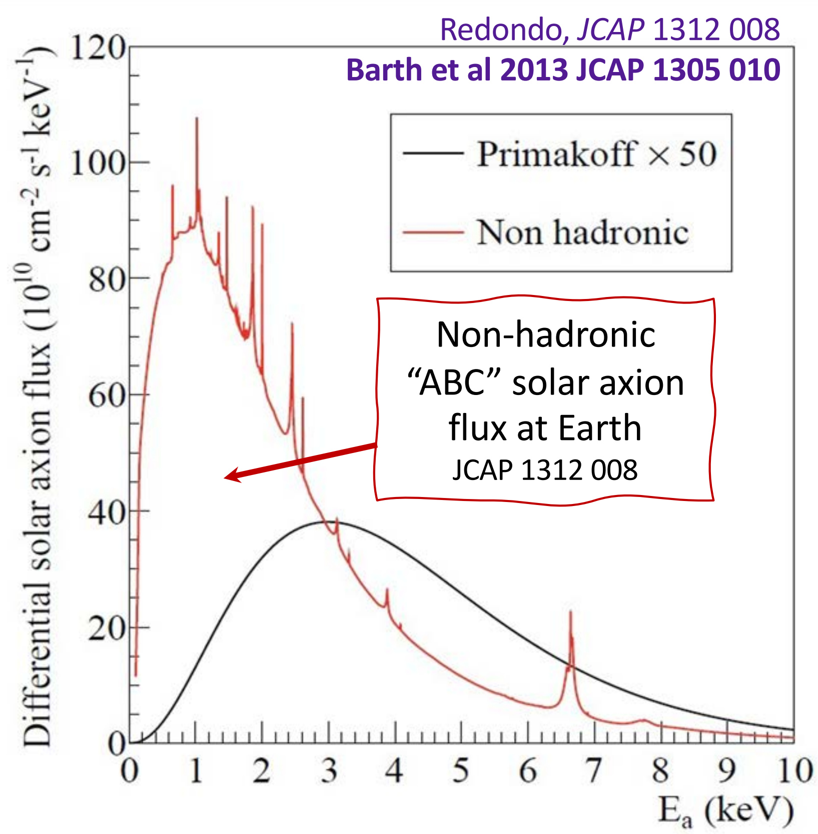
- Experimental conditions for BabyIAXO:
  - P_{max}(helium-4) $\simeq$ 1bar
  - T(average) $\simeq$ 295K

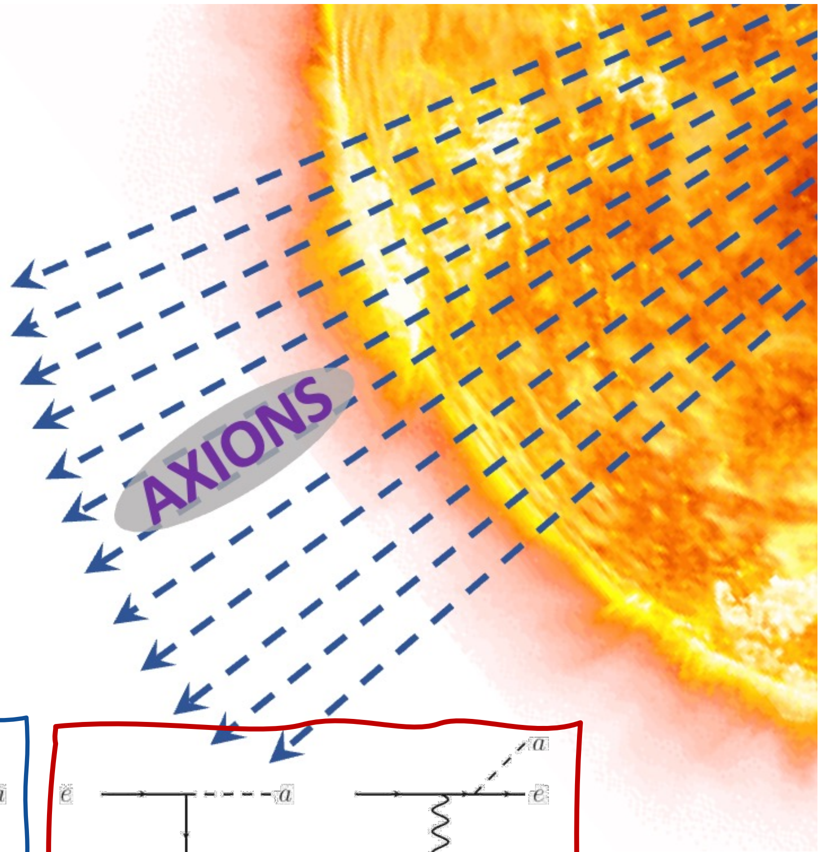
BabyIAXO



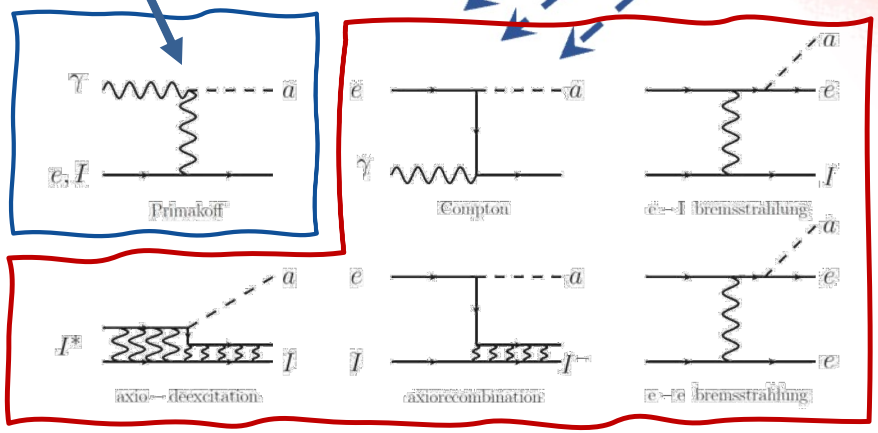
Solar Axion Searches

- Additionally to Primakoff: "ABC axions" which m be ✕100 more intense but model-dependent





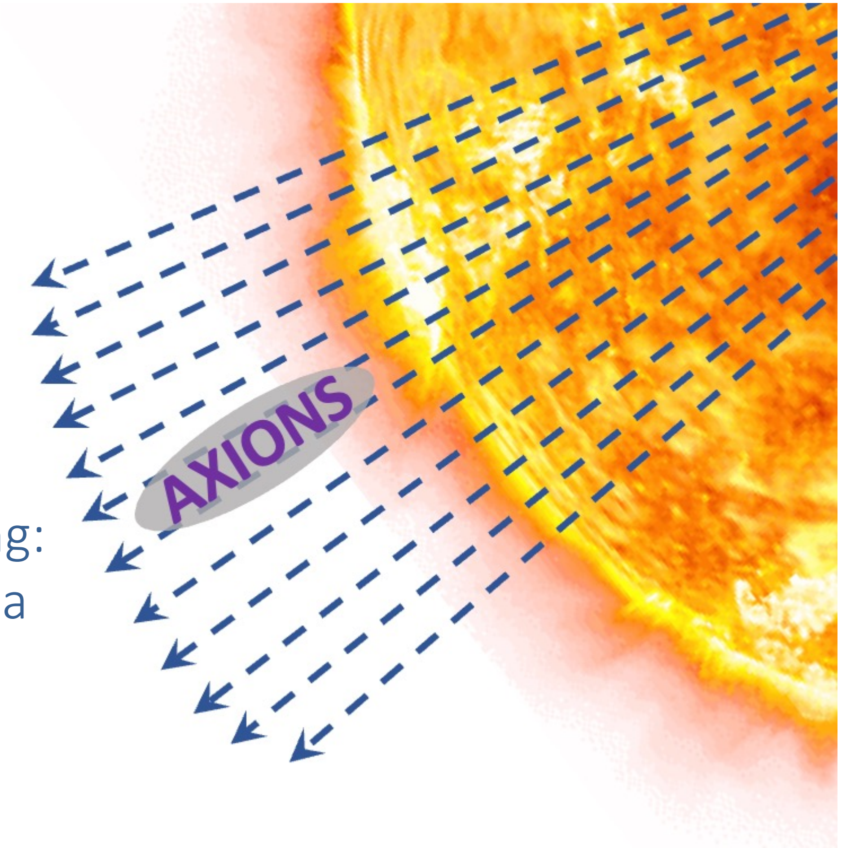
Primakoff axions

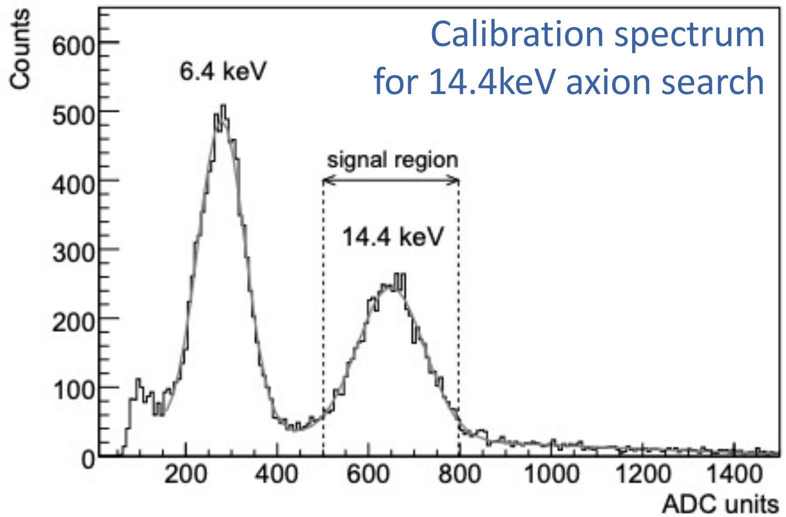


Non-hadronic or "ABC" axions (if the axion couples to electrons (g_{ae}))

Solar Axion Searches

- Via axion-nucleon couplings can also observe monochromatic lines from nuclear transitions
  - keV axions emitted in the M1 transition of Fe-57 nuclei (14.4 keV) and Tm-169 (8.4keV)
  - MeV axions from ^{7}Li (0.478 MeV) and $D(p;g)^{3}He$ (5.5 MeV)
- Axions-nucleon coupling g_{aN} especially intriguing: if the axion has couples via g_{aN} , it is most likely a QCD axion





Calibration spectrum for 14.4keV axion search

$$\Phi_a = 5.06 \times 10^{23} \, (g_{aN}^{\text{eff}})^2 \, \text{cm}^{-2}\text{s}^{-1}$$

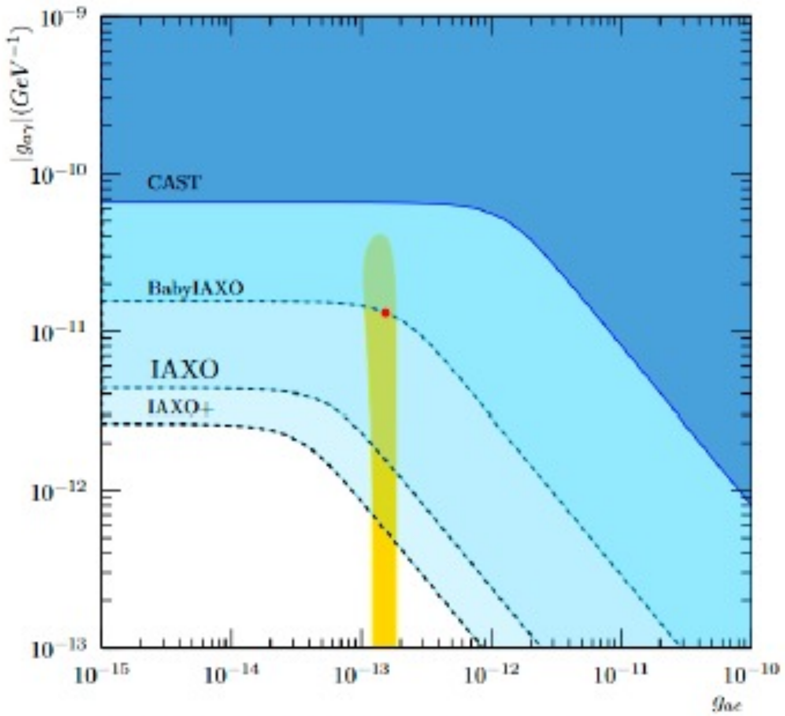
Di Luzio *et al* 2022 *Eur. Phys. J.* C 82:120
CAST collaboration *et al* 2009 *JCAP* 12 002
D. Miller *et al* 2010 JCAP 1003 032
Derbin *et al 2023 Jetp Lett.* **118**, 160

Solar Axion Searches

Detection Helioscope

Parameter space showing the sensitivity of the experiments in the $g_{a\gamma}$ - g_{ae} plane

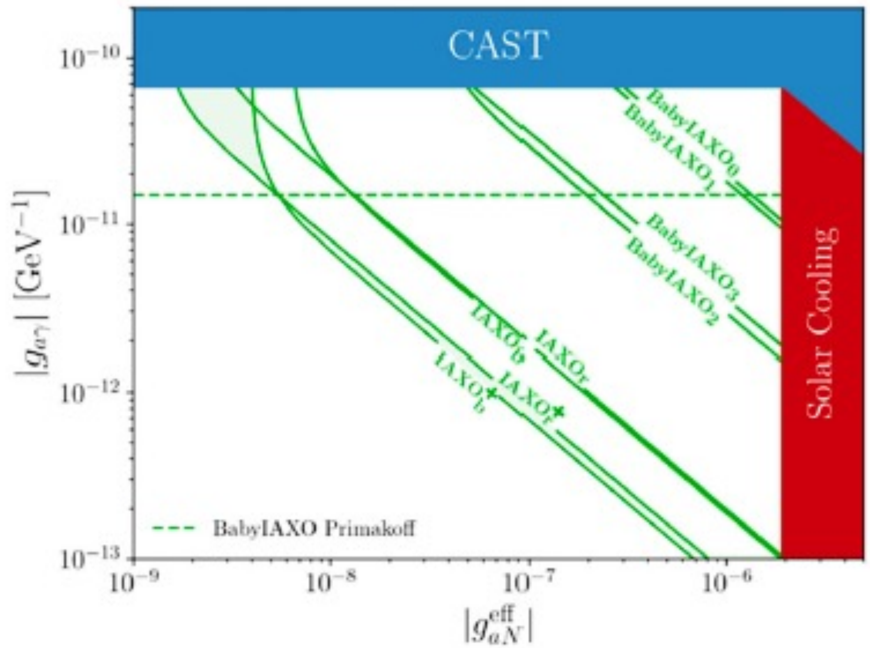
- Axion mass $m_a \simeq 1meV$



Physics potential of the International Axion Observatory (IAXO) JCAP06(2019)047

Parameter space showing the sensitivity of the experiments in the $g_{a\gamma}$ - g_{aN} plane

- Axion mass $m_a \simeq 20meV$

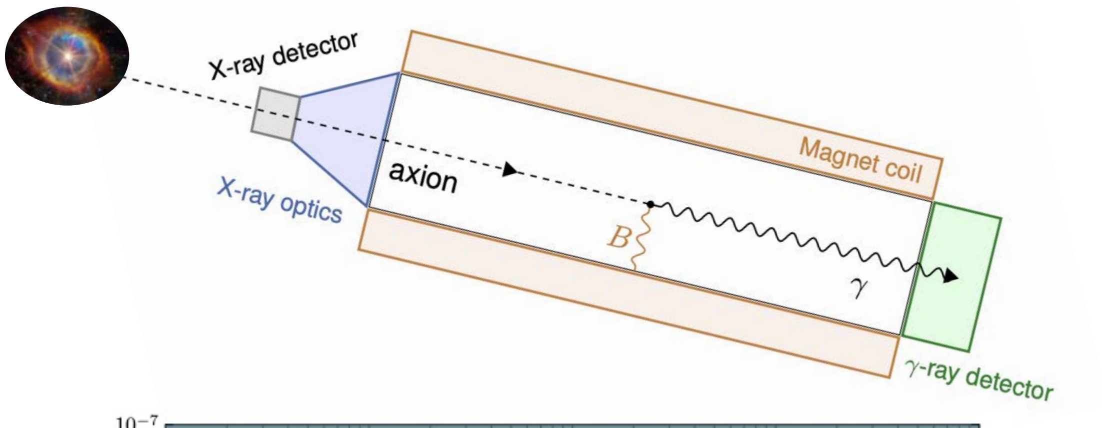


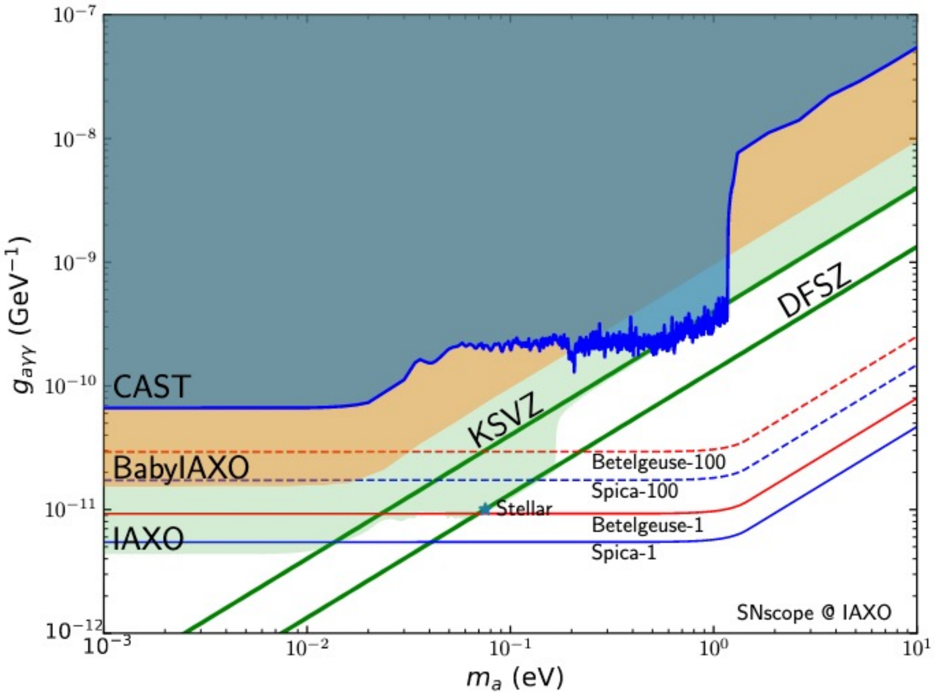
Probing the axion–nucleon coupling with the next generation of axion helioscopes Eur. Phys. J. C (2022) 82:120

Axion from galactic supernova

- If a sufficiently close-by galactic SN explodes, SN axions could be detectable at (Baby)IAXO.
- SN axions have O(100MeV) energies
- Requires IAXO to be equipped with large HE g-ray detector, covering all magnet bore, sufficient pointing accuracy, alert system in place
- Can be implemented complementary to baseline BabyIAXO setup by using opposite side of magnet.

Ge et al. 2008.03924



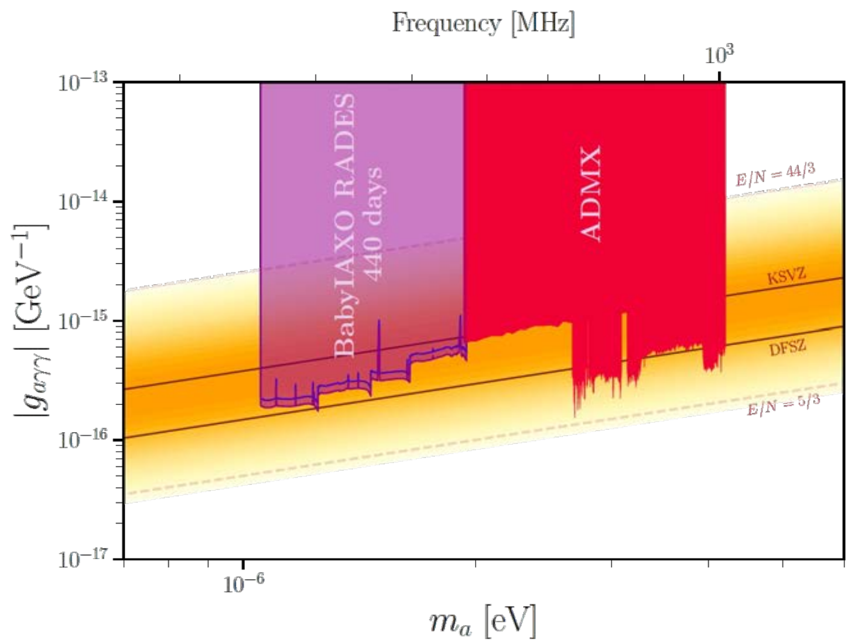


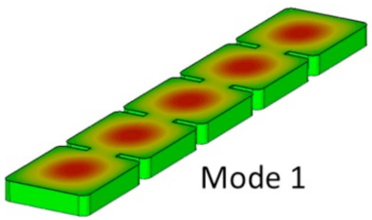
IAXO Physics

Beyond baseline physics

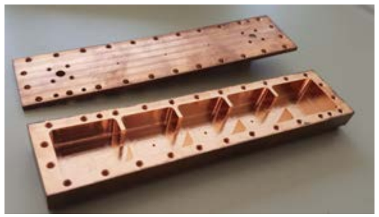
- Use of (Baby)IAXO large magnetic volume for axion DM setups
- Very competitive prospects for 1-2 meV axion searches.
- **DarkQuantum!**

Ahyoune et al. (RADES Collaboration) arxiv:2306.17243





New geometry concepts to scale in V but keeping high resonant *f*



JCAP 05 (2018) 040
JHEP 07 (2020) 084

I. Irastorza (U. Zaragoza), T. Kontos (École Normale Supérieure de Paris), S. Paraoanu (Aalto University), W. Wernsdorfer (KIT)

ERC-Synergy Grant DarkQuantum obtained !

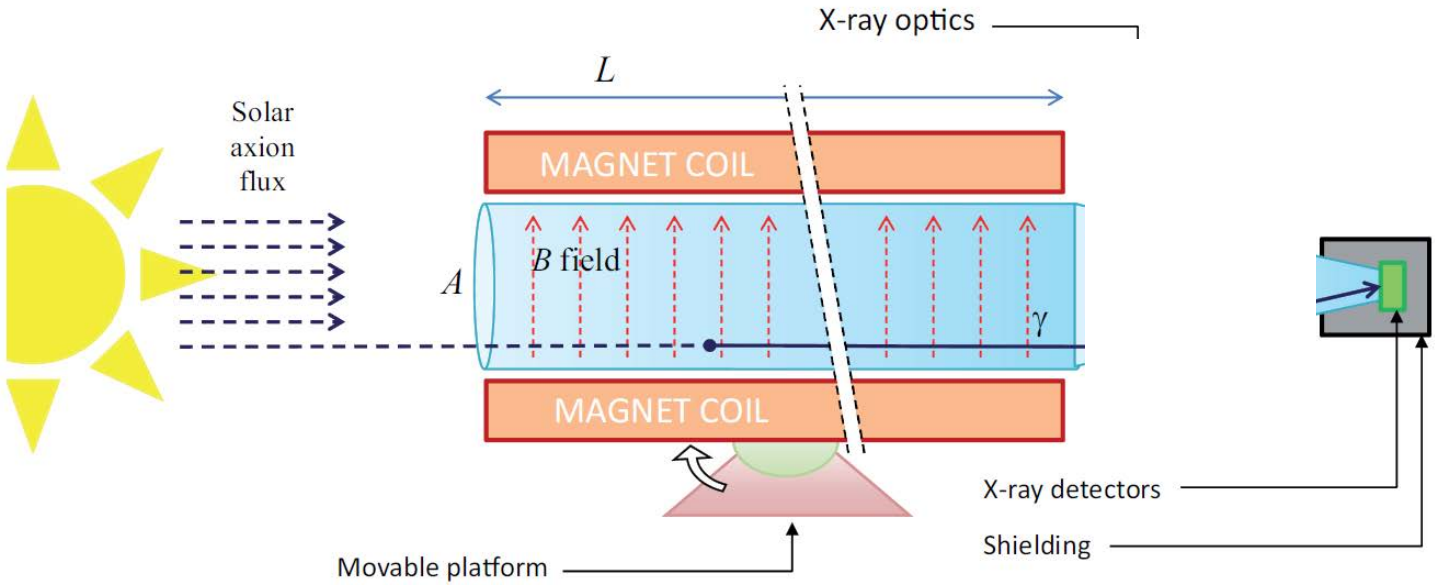
- Develop quantum tech for axions
- Quantum-enhanced haloscope in BabyIAXO
- Connection with experts (cryo, quantum,...)
- Contribution to magnet

DarkQuantum

How do we surpass the current state-of-the-art results?

Helioscope Figure of Merit

Enhanced axion helioscope:
Irastorza et al 2011
JCAP 1106, 013



Measure of sensitivity to axion-photon interaction:

$$g^4_{a\gamma} \propto \underbrace{(BL)^{-2} A^{-1}}_{\text{magnet}} \times \underbrace{t^{-1/2}}_{\text{exposure}} \quad \underbrace{b^{1/2}\varepsilon^{-1}}_{\text{detectors}}$$

magnet
- B = magnetic field
- L = magnet length
- A = cross-sectional area

exposure
- t = time

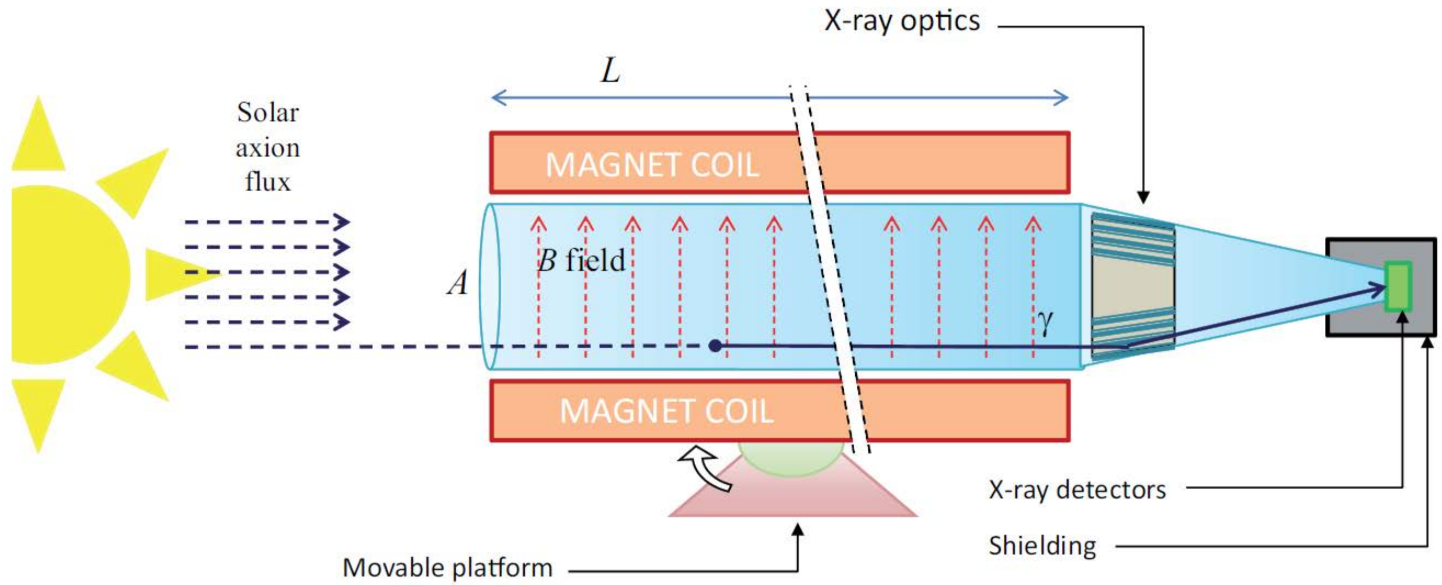
detectors
- b = background
- ε = efficiency

The smaller g_{ag} the better!

Solar Axion Searches

Helioscope Figure of Merit

Enhanced axion helioscope: Irastorza et al 2011 JCAP 1106, 013



Measure of sensitivity to axion-photon interaction:

$$g_{a\gamma}^{4} \propto (BL)^{-2} A^{-1} \times t^{-1/2} \times s^{1/2}\varepsilon_0^{-1} \times b^{1/2}\varepsilon^{-1}$$

magnet
- B = magnetic field
- L = magnet length
- A = cross-sectional area

exposure
- t = time

optics
- s = spot size
- ε_0 = efficiency

detectors
- b = background
- ε = efficiency

The smaller g_{ag} the better!

Expect improvement for next gen (International Axion Observatory):
1–1.5 orders of magnitude in sensitivity to $g_{a\gamma}$ (factor of 10000-20000 in S/N)

Solar Axion Searches

Next-gen experiments

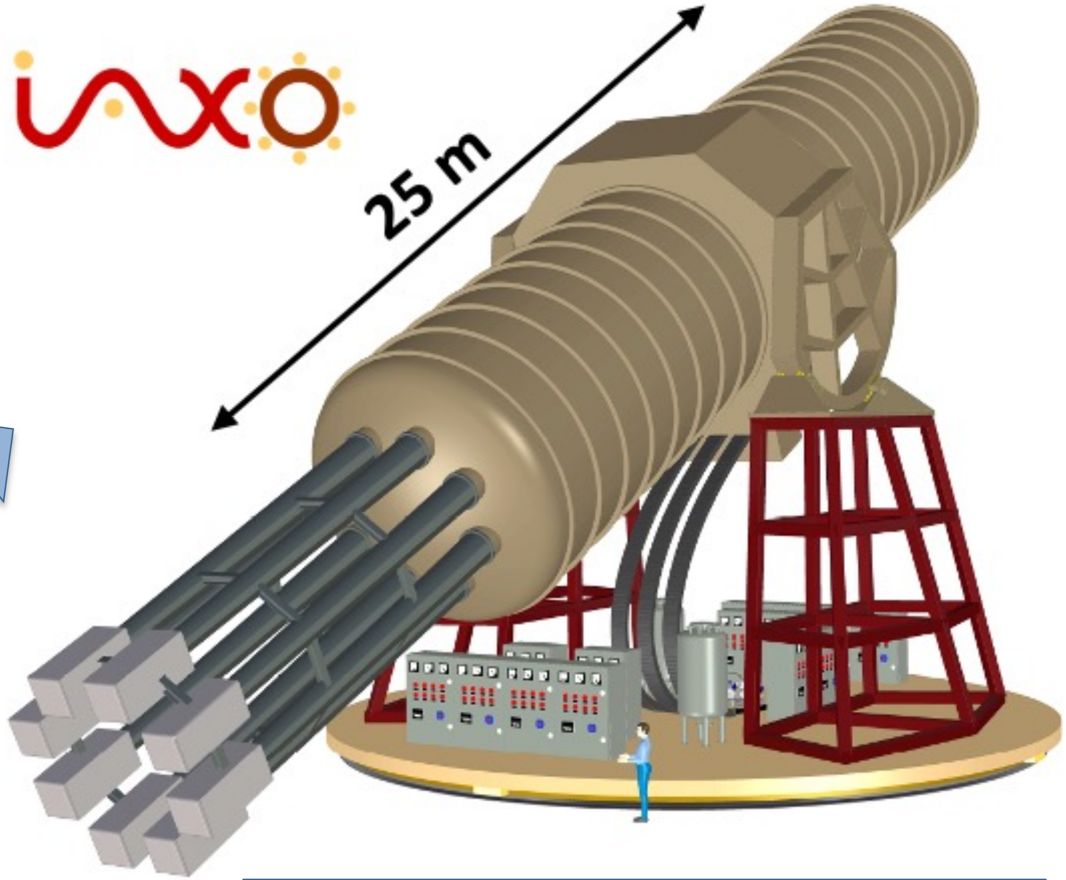
CAST

$g_{a\gamma} \lesssim 0.66 \times 10^{-10}$ GeV^{-1}

Baby IAXO

$g_{a\gamma} \lesssim$ few 10^{-11} GeV^{-1} (expected)

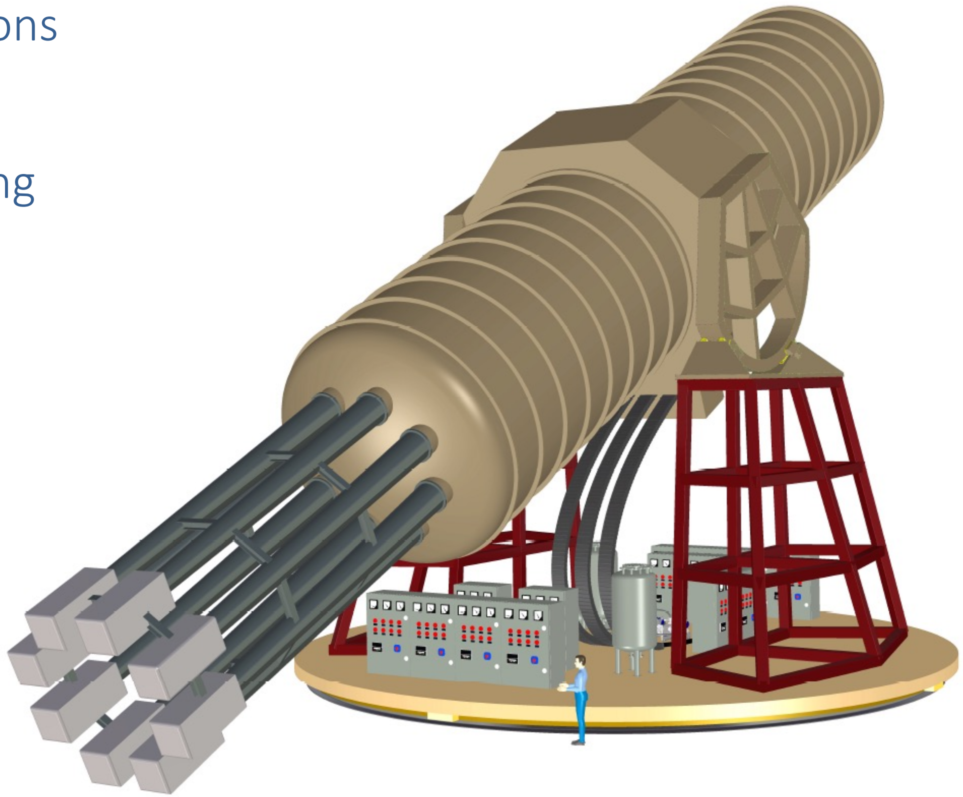
IAXO



$g_{a\gamma} \lesssim$ few 10^{-12} GeV^{-1} (expected)

Solar Axion Searches

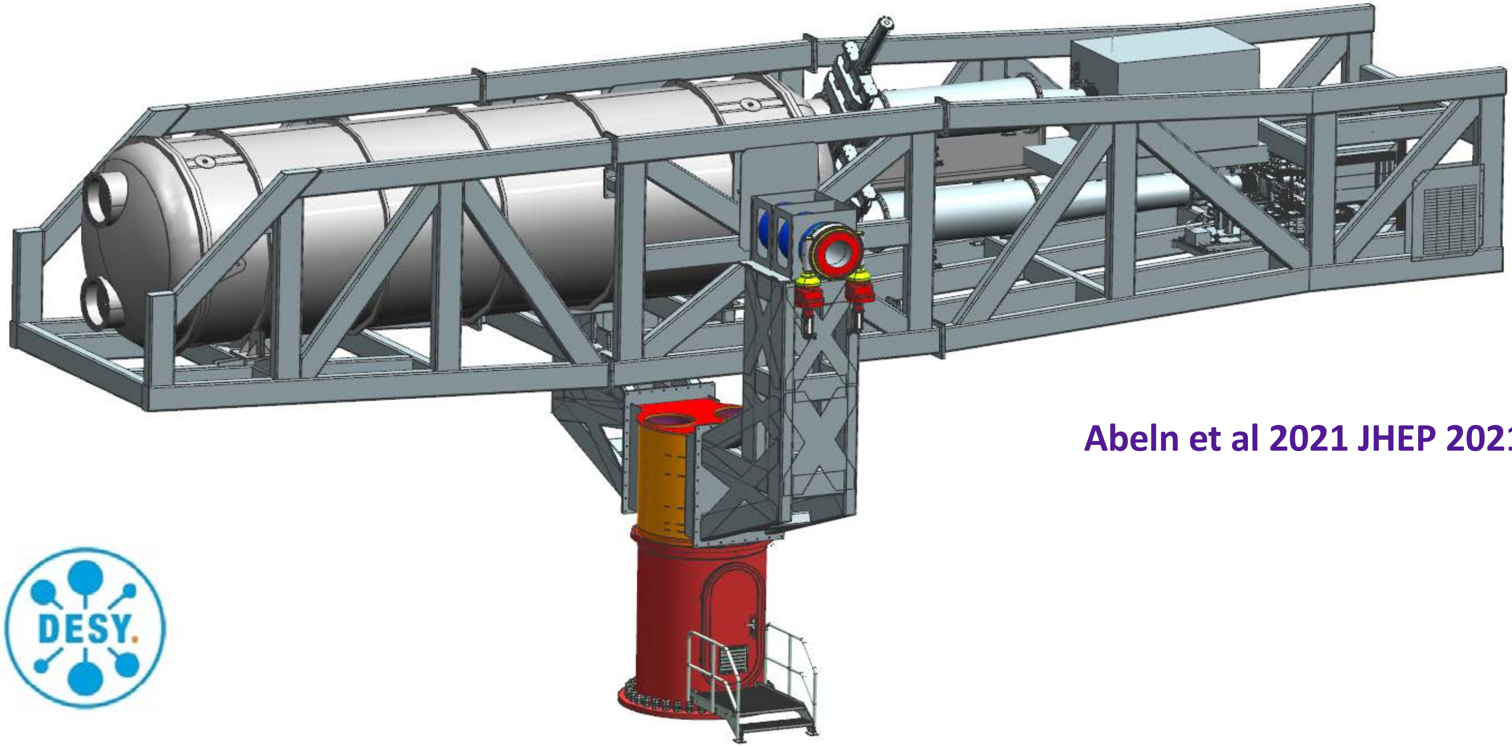
- **INTERNATIONAL AXION OBSERVATORY (IAXO)**
  - ✓ Next generation helioscope for solar axions
  - ✓ Mature and state-of-the-art technology
  - ✓ Purpose-built large-scale superconducting magnet
    - Toroidal geometry
    - 25 meters long, up to 5.4 T.
    - >300 times larger FoM than CAST magnet
    - 8 conversion bores of 60 cm Ø
  - ✓ 8 detection lines
    - X-ray optics with 0.2 cm² focal spot
    - Ultra-low bgrd detectors
  - ✓ 50% of Sun-tracking time.



Armengaud et al 2014 JINST 9 T05002
Irastorza et al 2011 JCAP 1106, 013

Solar Axion Searches

- **BABYIAXO =INTERMEDIATE EXPERIMENTAL STAGE BEFORE IAXO**
  - ✓ Technological prototype of IAXO with only two magnet bores (10 m, Ø 70 cm) to be installed at DESY
  - ✓ Relevant physical outcome (~10✕ CAST B^2L^2A)
  - ✓ Magnet will be upscalable version for IAXO
  - ✓ X-ray optics/detectors close to final IAXO configuration (focal length, performance)



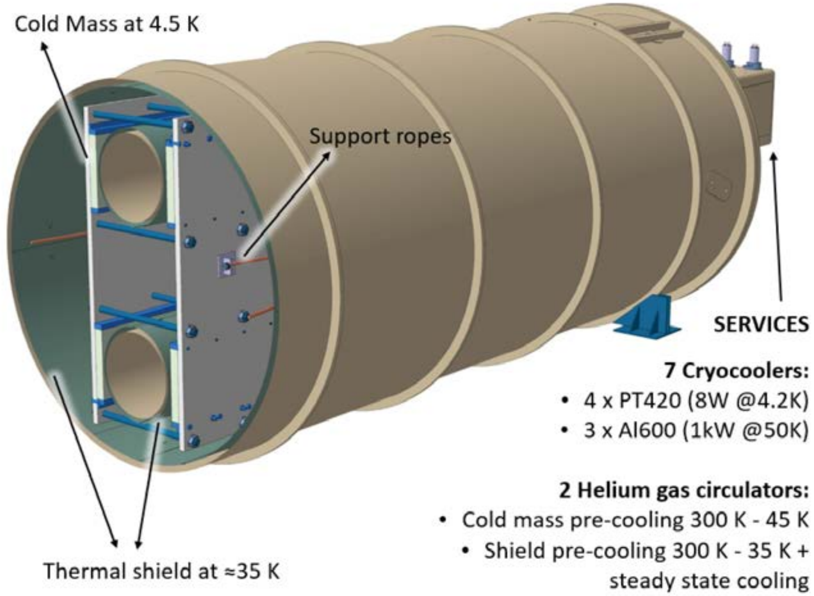
Abeln et al 2021 JHEP 2021 137

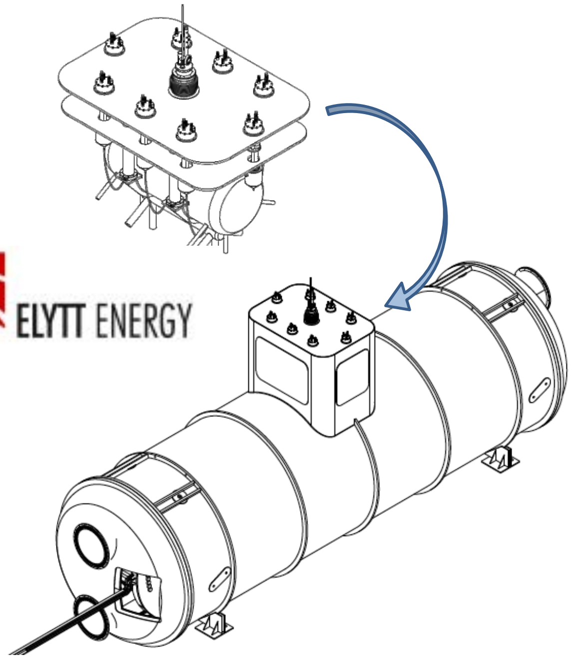
Solar Axion Searches

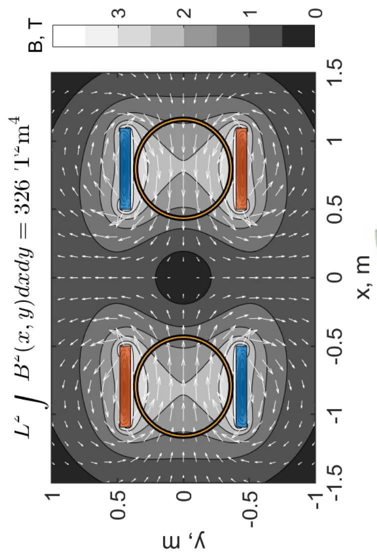
Solar Axion Searches

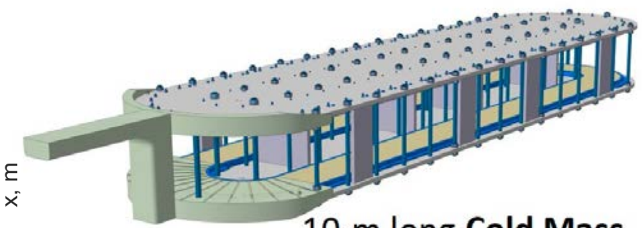
Solar Axion Searches

- ✓ BabyIAXO magnet to be operated at T ≤ 5 K featuring Nb-Ti-based superconducting coils with about 2 T in the bore
  - Nb-Ti is most affordable superconductor
  - It is also mechanically ductile and robust
  - Well studied work-horse conductor for most existing superconducting magnets





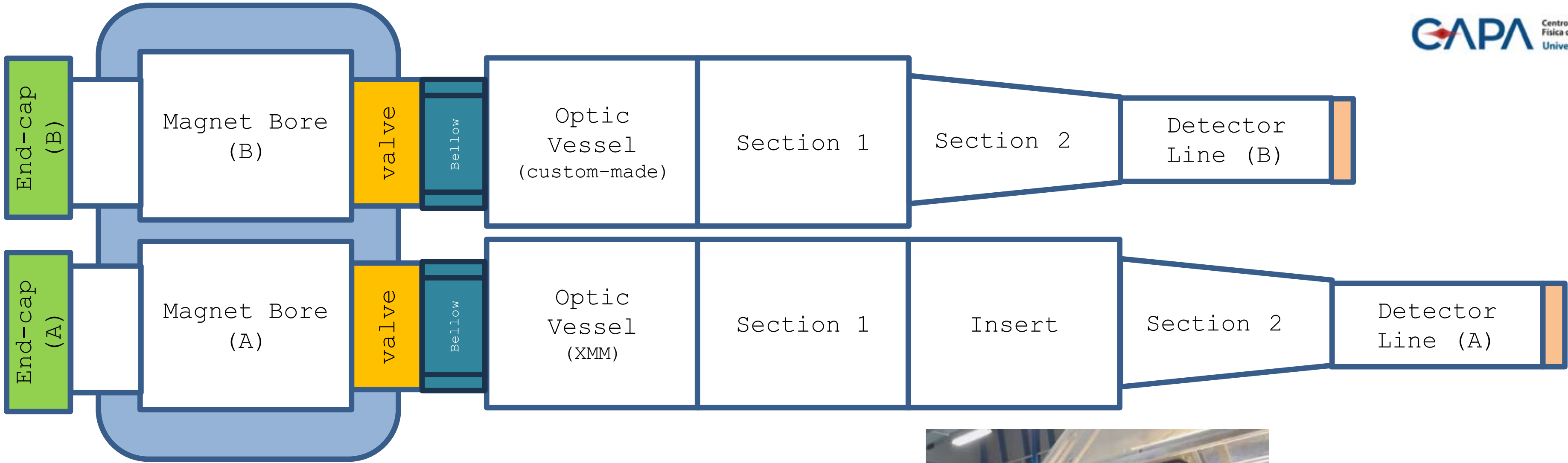


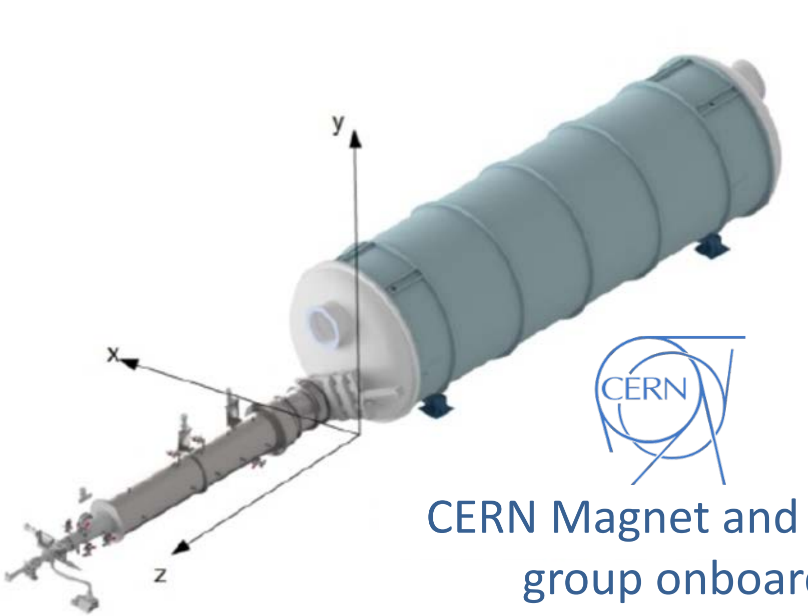


10-m long **Cold Mass**

MAGNET DEVELOPMENT

Solar Axion Searches





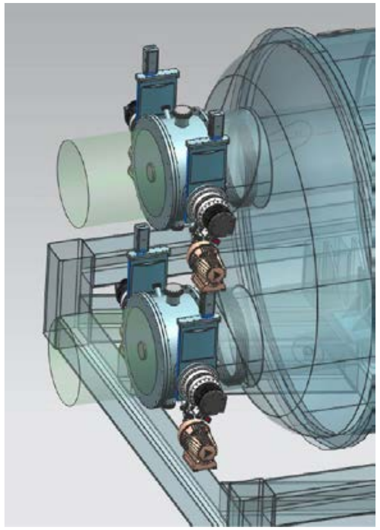
CERN Magnet and Cable group onboard

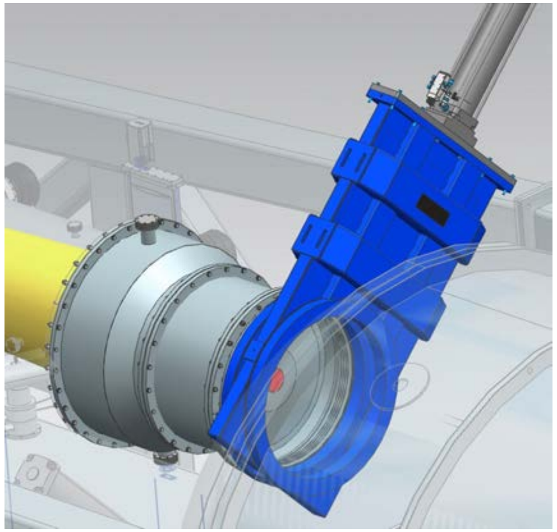
Contract for cable development recently signed

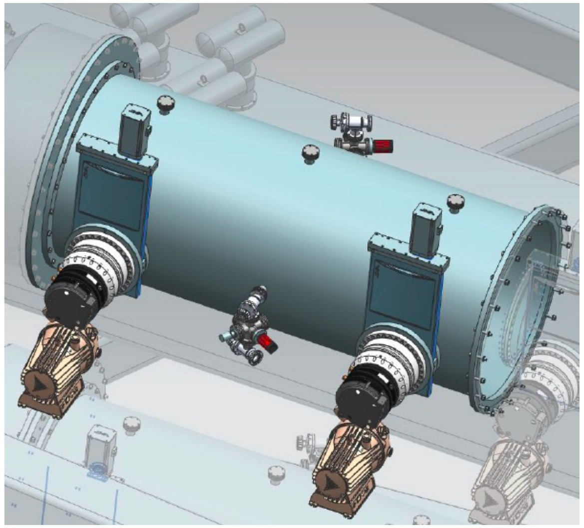
Ordering 2km of cable strand for winding tests

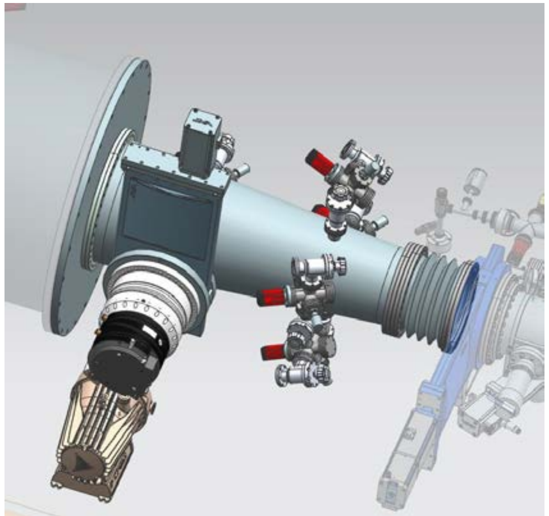
BEAM LINE

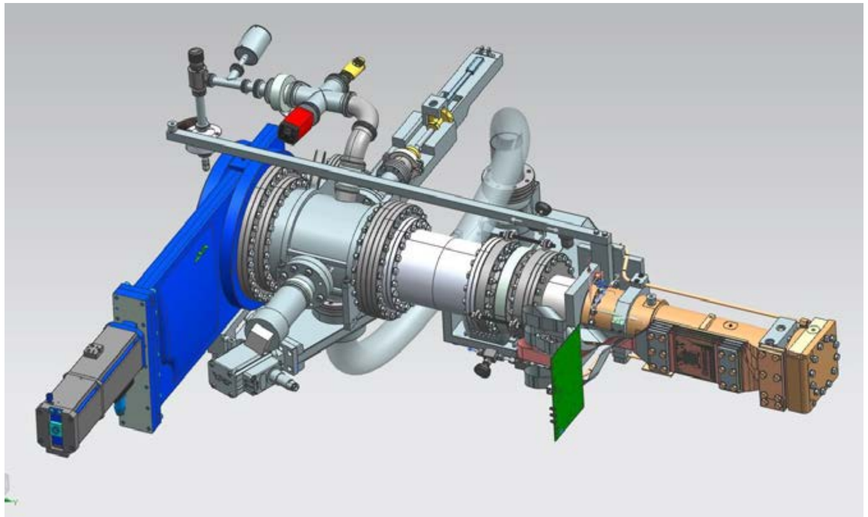
Solar Axion Searches


Magnet End-cap


Magnet Telescope Vessel


Beam line Section 1


Beam line Section 2


Detector line

BEAM LINE

Solar Axion Searches

Baby IAXO

- **BabyIAXO OPTIC APPROACH:**

Henriksen et al 2021 AO 60, 22
Irastorza et al 2015 JCAP 12, 008

✓ NEED: Maximized throughput efficiency (40-60%),
Small focal spot area (r < 2.5 mm)

✓ Baseline option (1-10 keV) (prototyping and R&D)
- Bore 1: Existing XMM flight-spare telescope (replicated optics)
- Bore 2: Custom IAXO optic (single- or multilayer-coated, segmented-glass or Al-foil Wolter-I, NuSTAR/XRISM/ATHENA)

✓ Beyond baseline (funding request pending)
- Lower threshold of 0.3 keV or better
- Add sensitivity at 14.4 keV

MAX PLANCK INSTITUTE FOR EXTRATERRESTRIAL PHYSICS

esa

Leveraging decades of research by NASA and ESA for space instrumentation: minimal risk and superior performance

Solar Axion Searches

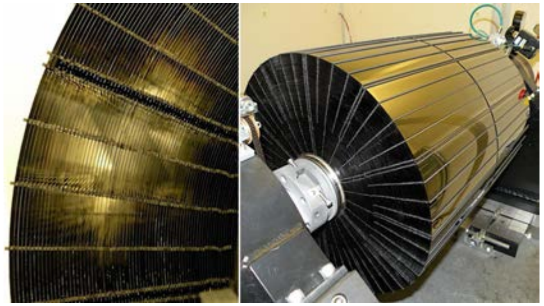
Baby IAXO

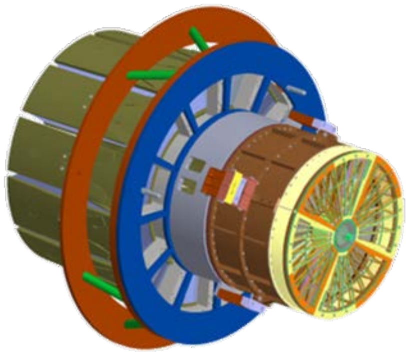
- ✓ IAXO-NuSTAR PATHFINDER @ CAST
- ✓ IAXO-CSGO/ATHENA PATHFINDER @ PANTER
- ✓ R&D for coating flat and curved substrates

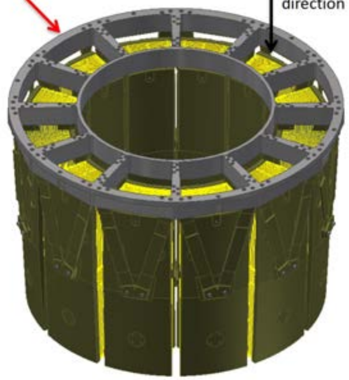
Massachusetts Institute of Technology

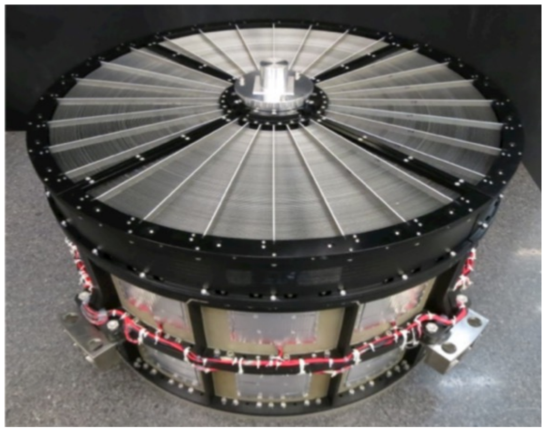
COLUMBIA UNIVERSITY IN THE CITY OF NEW YORK

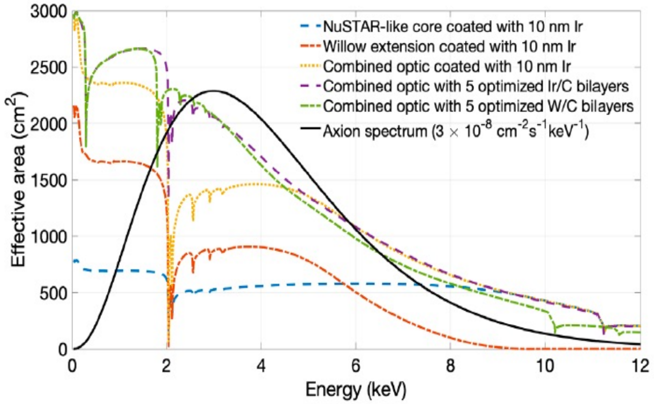
INAF ISTITUTO NAZIONALE DI ASTROFISICA



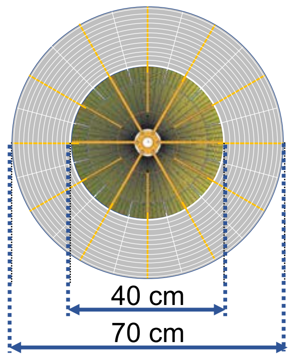








Custom optics
Hot and cold slumped glass



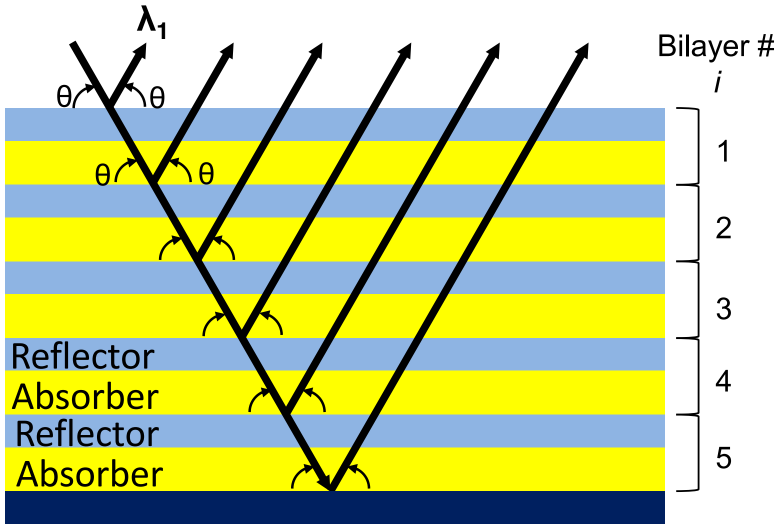
- CAD Design. Custom-optic beam line and cone of light.

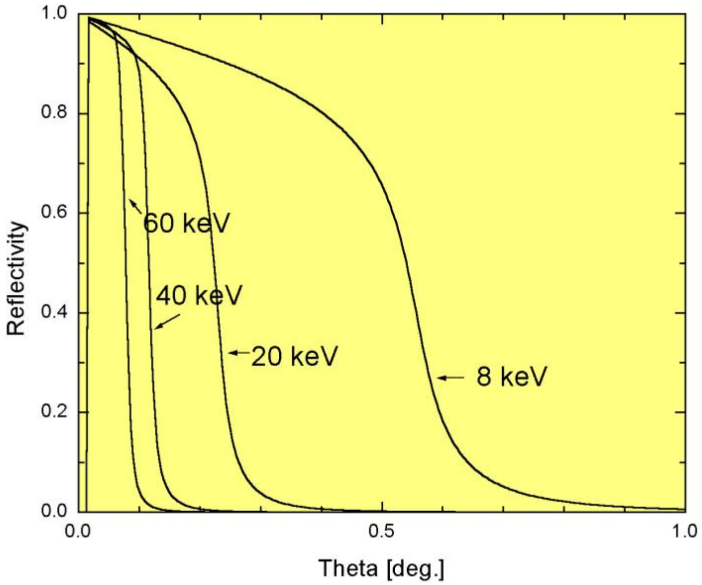
X-ray Optics

Solar Axion Searches

- **Intro to X-ray coatings:**
  - ✓ Reflectivity (from total external reflection) for metallic (Au, Pt, Ir) coatings high up to critical angle (θ_c) which depends on photon energy and mirror material
  - ✓ For larger x-ray energies (> 10-15 keV): metallic coatings no longer efficient enough (shallow angles, single reflection)
  - ✓ Multilayer coatings to extend energy range (or optimize response <10 keV) by making use of Bragg's law:

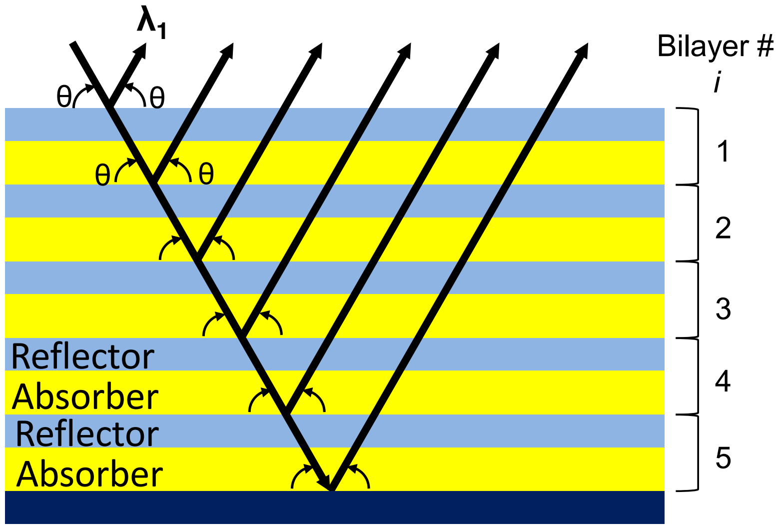
$$\lambda m = 2d\sin\theta_i$$

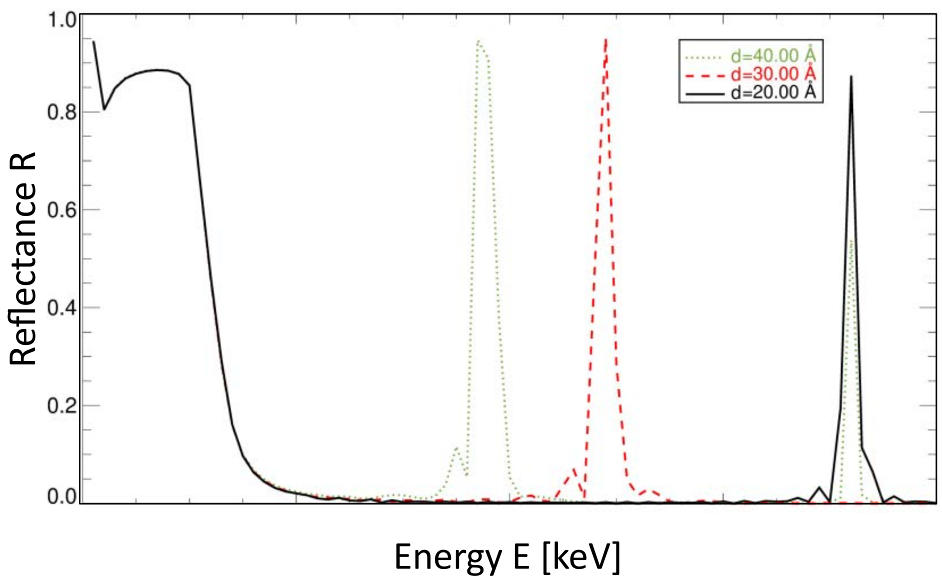




Solar Axion Searches

- **Intro to X-ray coatings:**
  - ✓ Reflectivity (from total external reflection) for metallic (Au, Pt, Ir) coatings high up to critical angle (θ_c) which depends on photon energy and mirror material
  - ✓ For larger x-ray energies (> 10-15 keV): metallic coatings no longer efficient enough (shallow angles, single reflection)
  - ✓ Tuning the period (with const-d throughout the stack) to maximize reflectivity for specific E
  - ✓ Larger energies require smaller periods—current practical limit is d = 1.5 nm

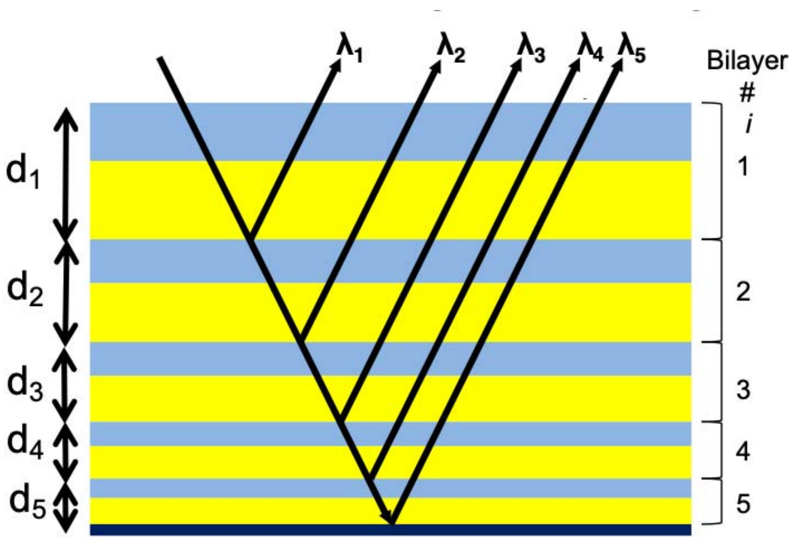


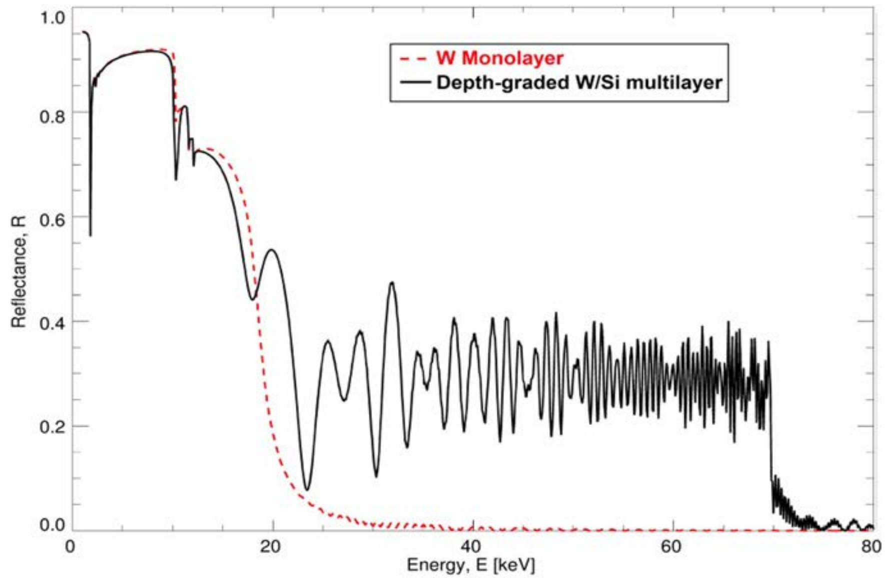


Solar Axion Searches

Baby IAXO

- **Intro to X-ray coatings:**
  - ✓ Reflectivity (from total external reflection) for metallic (Au, Pt, Ir) coatings high up to critical angle (θ_c) which depends on photon energy and mirror material
  - ✓ For larger x-ray energies (> 10-15 keV): metallic coatings no longer efficient enough (shallow angles, single reflection)
  - ✓ Other options like depth-graded d-spacing, aperiodic multilayers etc for increased flexibility, but extend energy width or tailored response is at the cost of high reflectivity



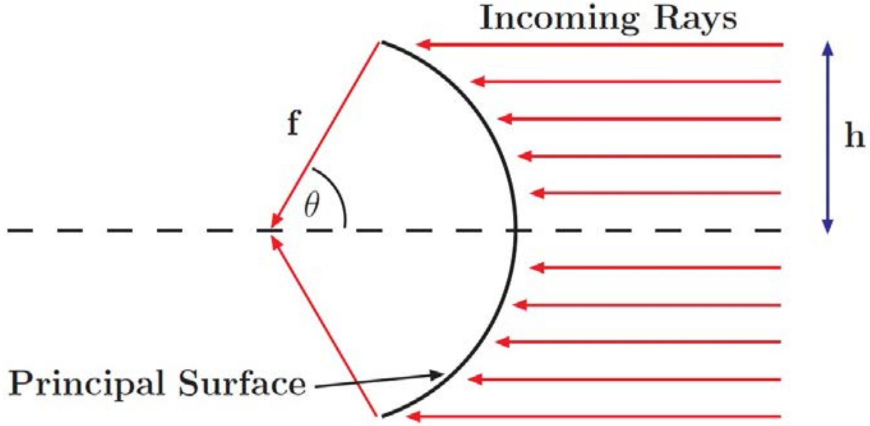


Multiparameter optimization that depends on science goals and priorities

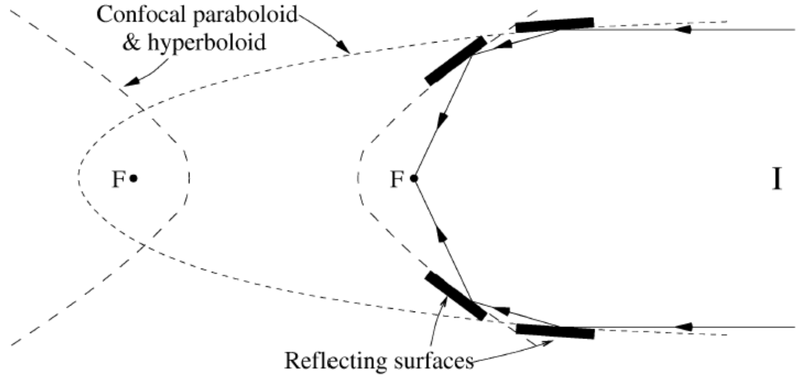
Solar Axion Searches

- **Intro to X-ray optics:**
  - ✓ X-ray optics are based on the principle of total external reflection (TER) below critical angle → Grazing incidence optics
  - ✓ In order to obtain sharp image (same focal spot for different h) over field of view, Abbe sine condition need to be satisfied:

$$f = \frac{h}{\sin(\theta)}$$



  - ✓ H. Wolter (1952): Two conic surfaces of revolutions to nearly satisfy Abbe sine rule
  - ✓ Three families of designs, one of which can be nested (Wolter I) and is widely used
  - ✓ Wolter I has properties similar to a thin lens

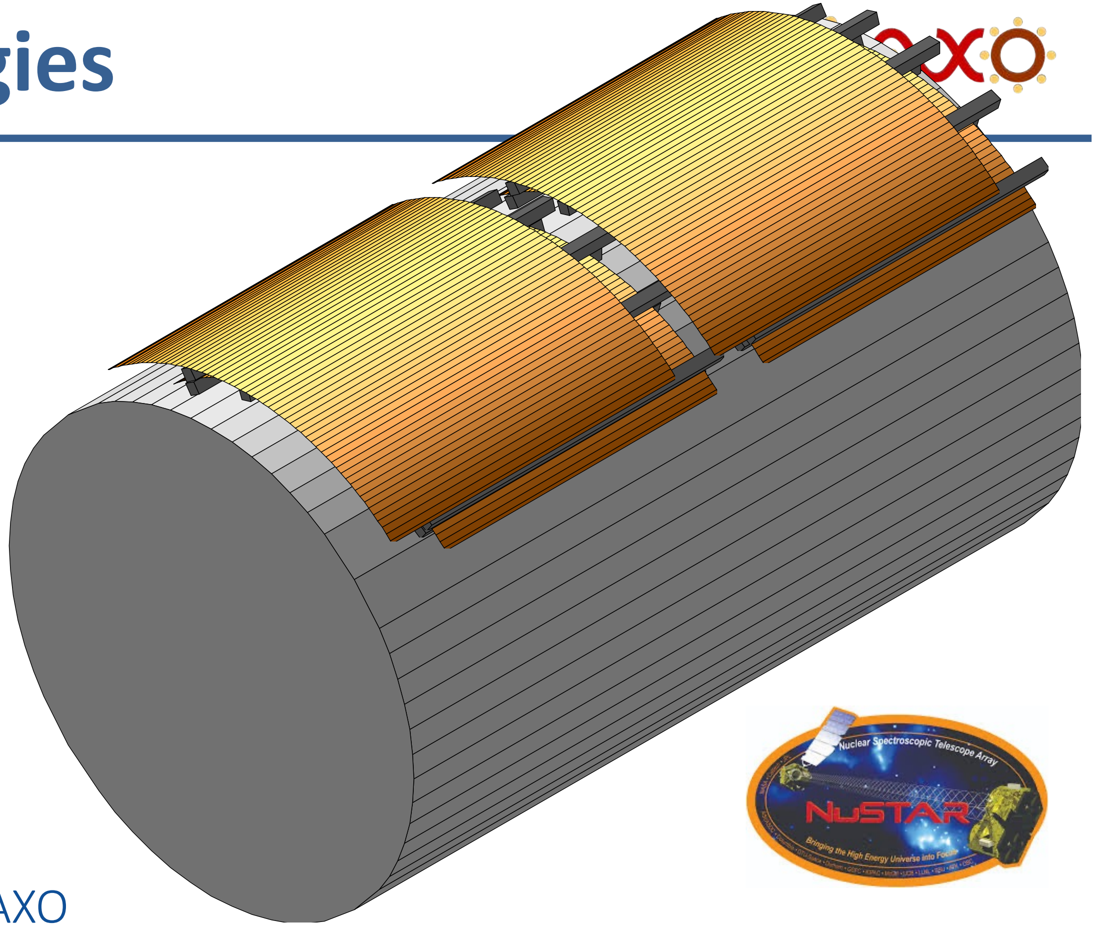


H. Wolter 1952 *Phys. Ann.* **6** 94.

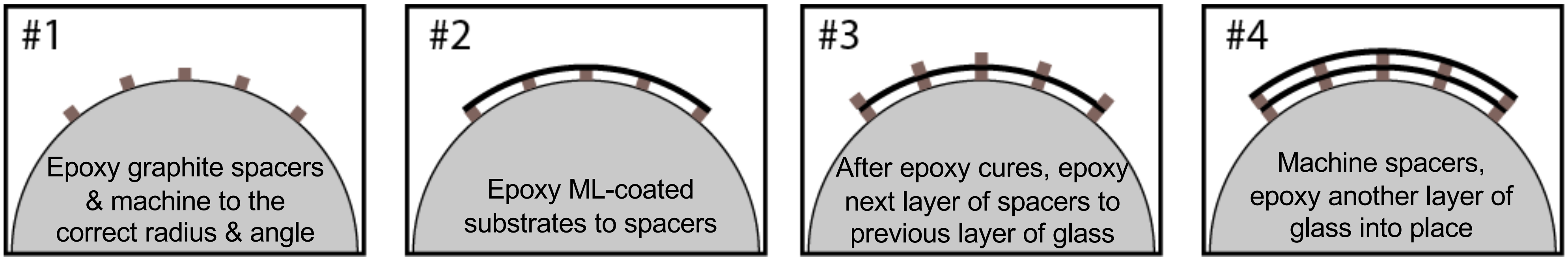
Advantages of Wolter optic:
Imaging capability, improvement of signal-to-noise, enables reduction of background

X-Ray optics technologies

- **Intro to X-ray optics:**
  - ✓ Segmented optics rely on several individual pieces of substrates to complete a single layer
  - ✓ Selected as baseline technology for (B)IAXO, because
    - Mature technology/Expertise
    - Single/multilayer coatings can be deposited
    - Cost-effective
    - Modest imaging requirements for IAXO



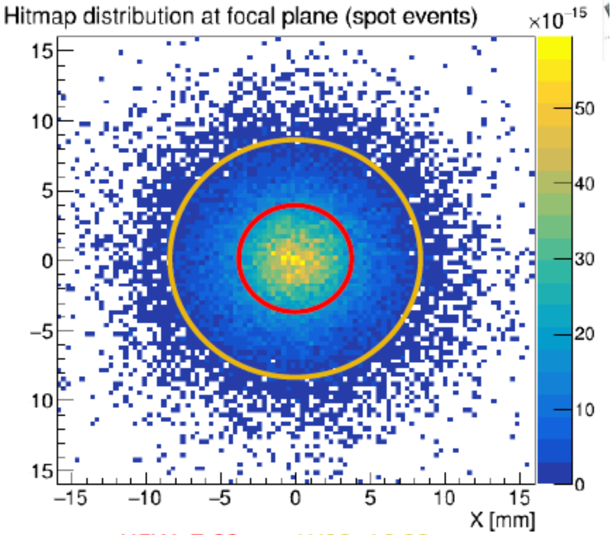
*JE Koglin et al. Proc. SPIE, **4851:607** (2003)*



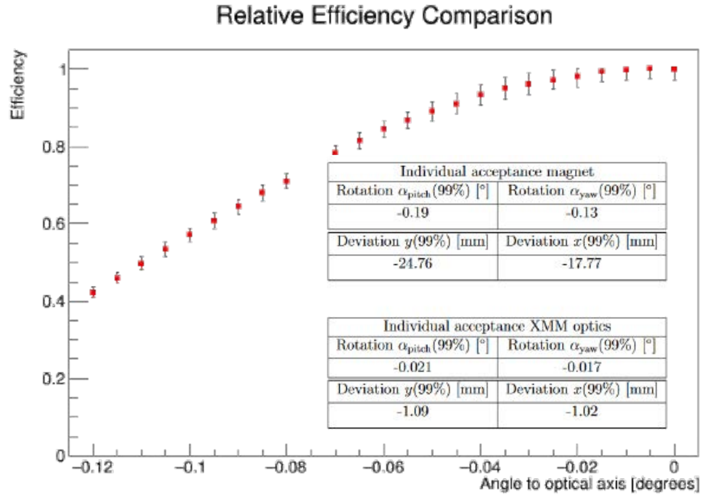
Solar Axion Searches

Baby IAXO

- ✓ Run with 100 000 events
- ✓ Primakoff flux
- ✓ Vacuum phase
- ✓ XMM optics
- ✓ No detector window
- ✓ REST v2.3.15

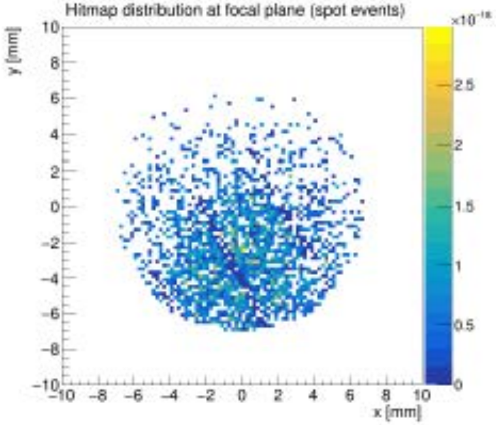


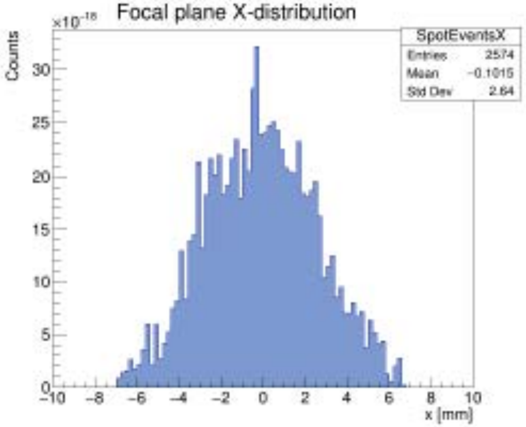
HEW=7.83mm W90=16.83mm
HEW=3.59′ W90=7.714′

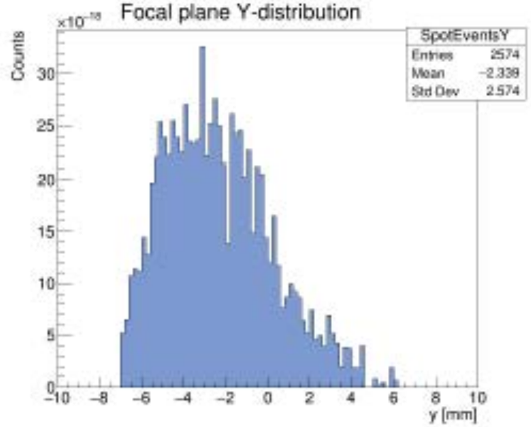


Magnet and optics rotation

Rare Event Searches Toolkit software



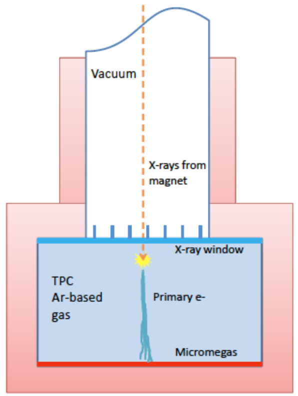


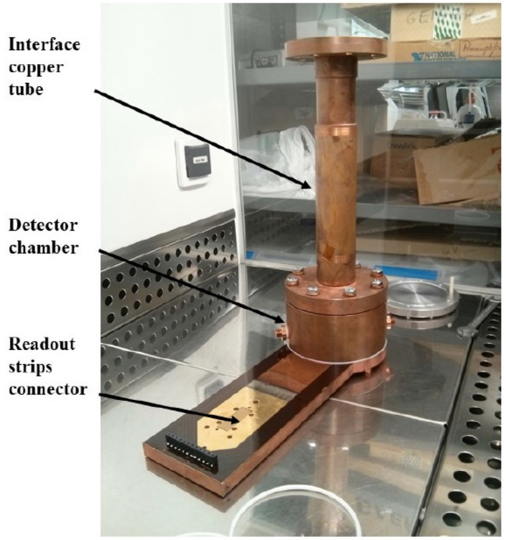


Acceptance Studies

Solar Axion Searches

- **NEED (Baseline 1-10 keV):**
  - ✓ Low background (<10^{-7} – 10^{-8} cts keV^{-1} cm^{-2} s^{-1})
    - Less than 1 event per 6 months of data taking!
    - Already demonstrated
      - ~8×10^{-7} c keV^{-1} cm^{-2} s^{-1} (in CAST 2014 result) and
      - 10^{-7} cts keV^{-1} cm^{-2} s^{-1} measured underground at the Canfranc Laboratory (LSC)
  - ✓ High detection efficiency
- **WANT:**
  - ✓ Low E-threshold (< 1 keV) and great E-resolution
    - Especially interesting for axion-electron measurements
    - Notably useful in case an axion signal is detected





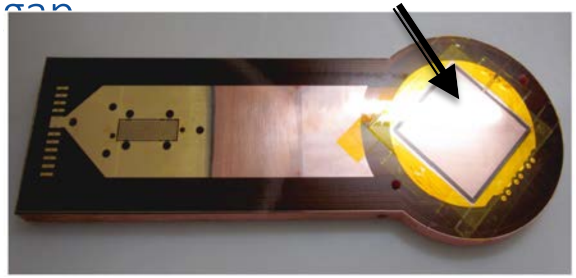
Detectors

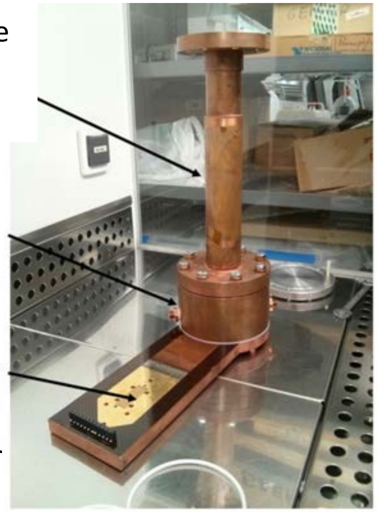
Micromegas best option to reach required low background Additional technologies considered and undergoing active R&D efforts (GridPix, MMC, TES, SDD)

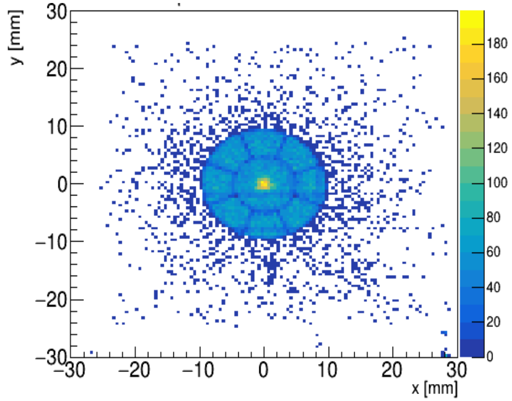
Solar Axion Searches

BabyIAXO

- **Microbulk Micromegas detectors**
- ✓ Very homogeneous amplification uniform gain
- ✓ Intrinsically radiopure
- ✓ Good energy and spatial resolution
- ✓ Pixelized readout gives topological information
- ✓ Signal reaches the active volume through a mylar window
- ✓ X-rays ionize the gas in the conversion region and the produced signal is read by the Micromegas
- ✓ Data is analyzed with the REST-for-Physics framework (github.com/rest-for-physics).







X-ray window

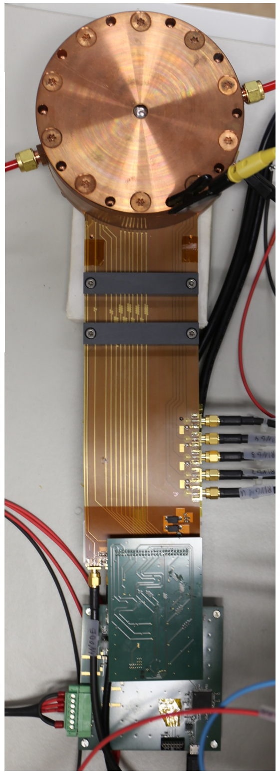
Rare Event Searches Toolkit software

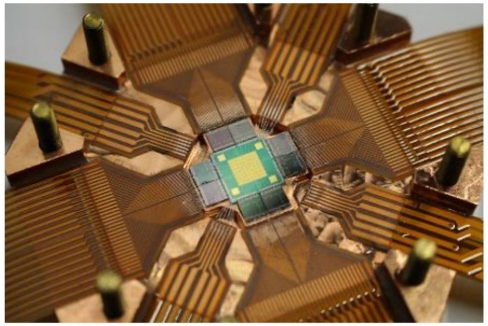
Detectors

Solar Axion Searches

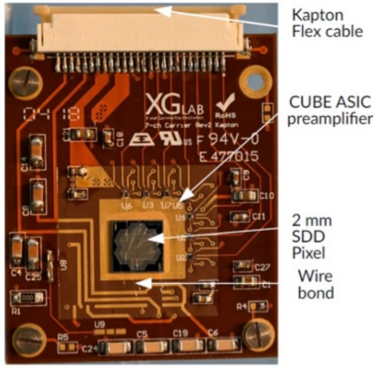
- Beyond baseline, "high-precision" detectors
- ✓ Better threshold & energy resolution
- ✓ Design and material optimization ongoing in all fronts
- ✓ Background studies with different shielding configurations
- ✓ DALPS project (French ANR)

ERC-StG (2020)
M.Meyer
To understand bkg in TES

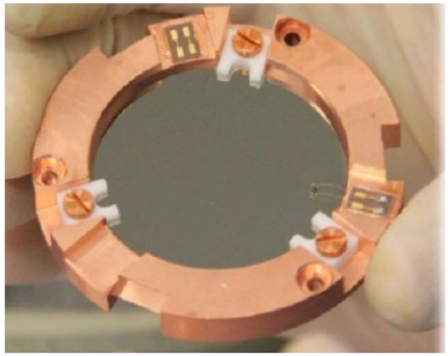




MMC: Metallic Magnetic calorimeters



SDD: Silicon Drift Detectors



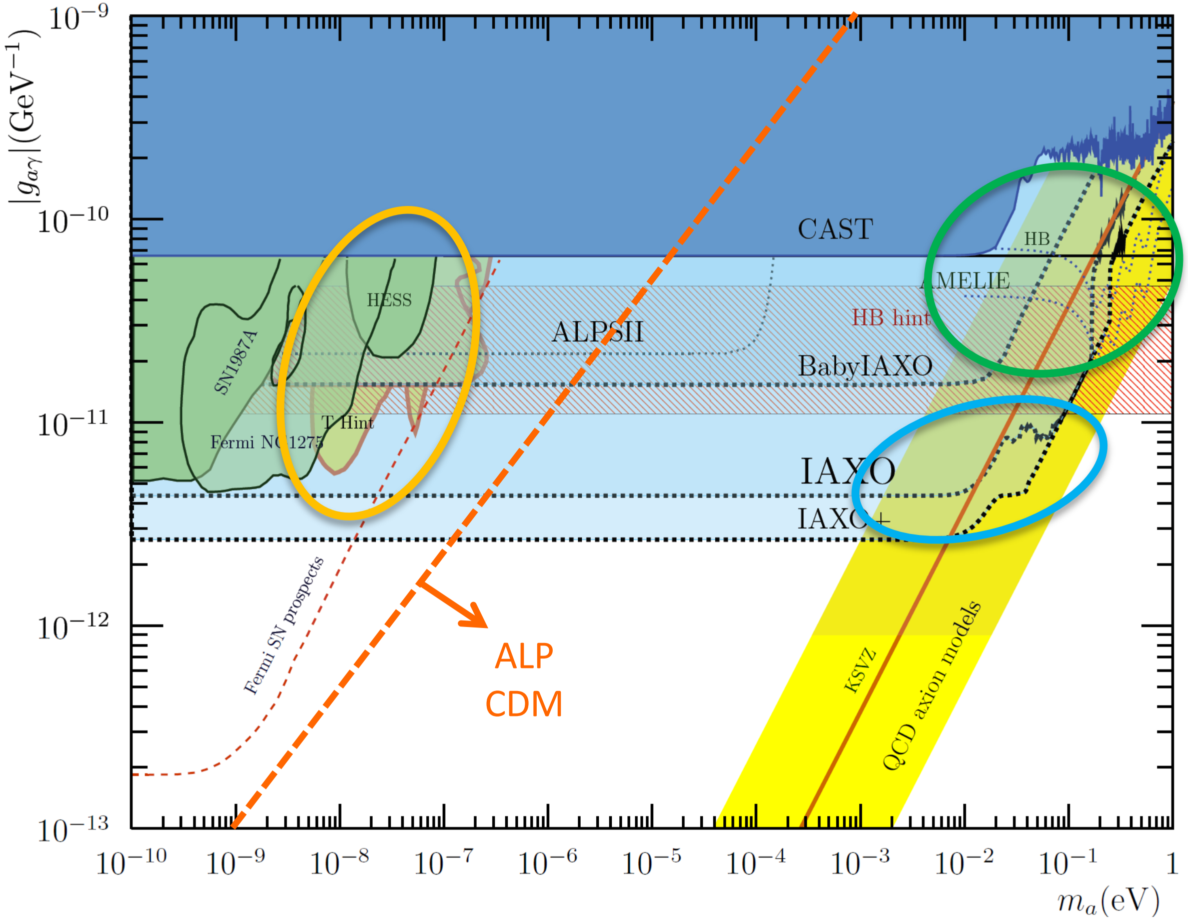
TES: Transition Edge senstors

Gridpix

Detectors

- ✓ Currently multiple new IAXO MM prototypes running in different locations (incl. Canfranc Underground Lab) with continuous improvements being made
- ✓ R&D ongoing for new detector technologies for high precision

Sensitivities



BabyIAXO prospects:
10 x $MFOM_{CAST}$ + optics and detector from conservative scenario of LoI

IAXO: > 300 x $MFOM_{CAST}$ +optics and detector improvements

IAXO+: Enhanced scenario with x 10 (x4) higher FOM (MFOM) with respect LoI

IAXO will probe large parts of QCD axion model space (KSVZ, DFSZ) including viable DM models

"ALP miracle" region: ALPs solving both DM & inflation (Daido et al. 2017 arXiv:1710.11107)

Large fraction of the axion & ALP models invoked in the "stellar cooling anomaly" (g_{ae} particularly interesting for this)

IAXO will fully explore ALP models invoked to solve the "transparency hint"

IAXO will also be able to probe large parameter space for CDM ALPs

Armengaud et al 2019, JCAP 1906, 047

Armengaud et al 2019, JCAP 1906, 047

IAXO as a generic axion(-like) detection facility

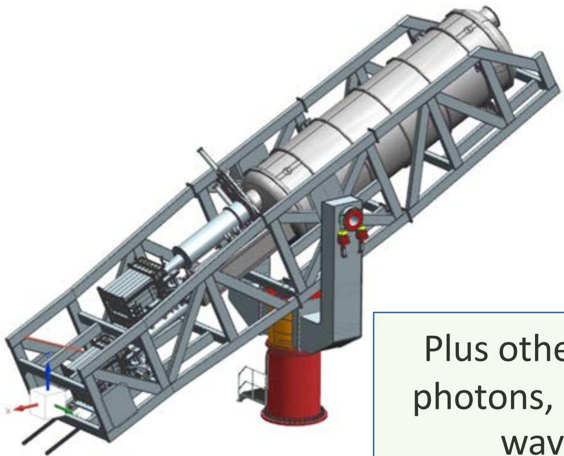
- (Baby)IAXO constitutes a great infrastructure that can be used to target other physics goals beyond Primakoff solar axions:

Other (non-Primakoff) solar axion production mechanisms (a-e, a-N, …)

Post-discovery "precision" physics (solar B and T, axion mass and model)

Axions from close-by supernova explosions

Dark Matter axions: haloscope setups inside the BabyIAXO bores

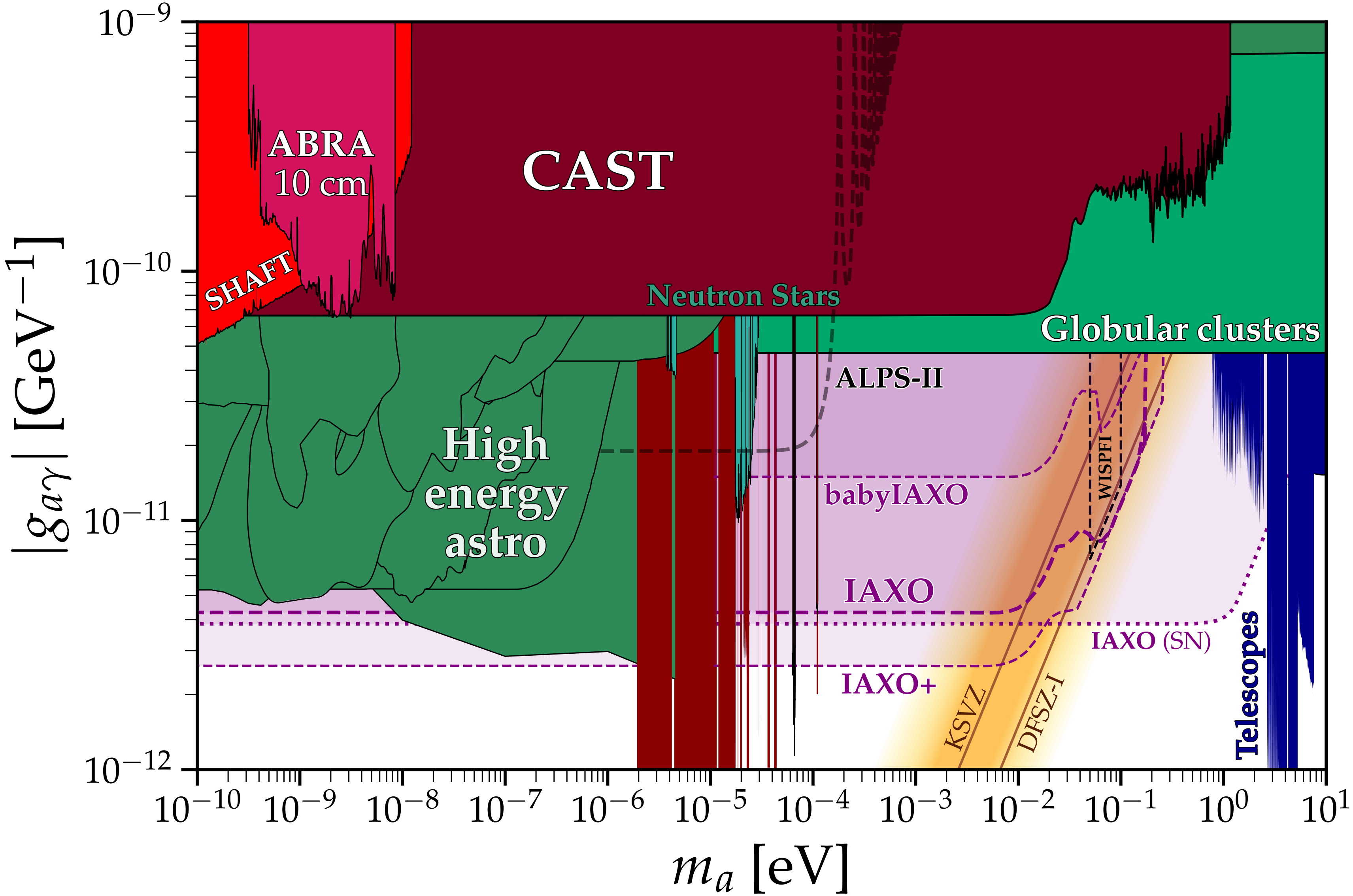


Plus other WISPs, such as hidden (dark) photons, chame-leons, etc., gravitational wave searches and NS studies

Beyond Base Line Science

Sensitivities

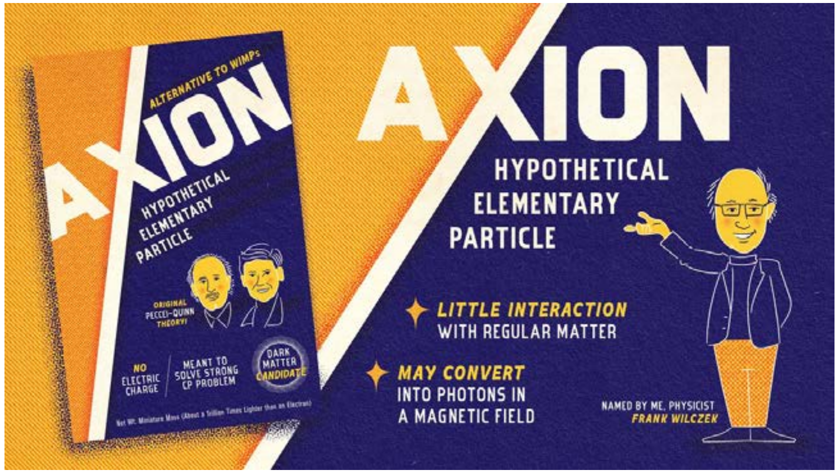
Adapted from https://github.com/cajohare/AxionLimits



Conclusions

- ✓ Axions are well motivated dark matter candidates simultaneously solving strong CP
- ✓ Axions (and axion-like particles) can be searched for in a variety of laboratory experiments: Haloscopes, Helioscopes and LSTW experiments
- ✓ Solar axion searches probe large regions of well-motivated axion parameter space
- ✓ BabyIAXO (IAXO) targets axion discovery with sensitivities down to a few 10^{-11} (10^{-12}) GeV^{-1} in $g_{a\gamma}$
- ✓ Intriguing IAXO physics cases beyond axion-photon (g_{ae}, g_{aN}, QCD, ALPs, astrophysical hints, dark photons, dark energy…)

THANK YOU FOR YOUR ATTENTION!



IAXO & BabyIAXO

Full members:

Kirchhoff Institute for Physics, Heidelberg U. (Germany) | IRFU-CEA (France) | CAPA-UNIZAR (Spain) | INAF-Brera (Italy) | CERN (Switzerland) | ICCUB-Barcelona (Spain) | Petersburg Nuclear Physics Institute (Russia, on hold) | Siegen University (Germany) | Barry University (USA) | Institute of Nuclear Research, Moscow (Russia , on hold) | University of Bonn (Germany) | DESY (Germany) | University of Mainz (Germany) | MIT (USA) | LLNL (USA) | University of Cape Town (S. Africa) | Moscow Institute of Physics and Technology (Russia , on hold) | Technical University Munich (TUM) (Germany) | CEFCA-Teruel (Spain) | U. Polytechnical of Cartagena (Spain) | U. of Hamburg (Germany) | MPE/PANTER (Germany)

Associate members:

DTU (Denmark) | U. Columbia (USA) | SOLEIL (France) | IJCLab (France) | LIST-CEA (France)