

CdTe and CdZnTe detectors for Hard X-ray imaging spectroscopy



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#### www.cea.fr

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### Intro

State of the art

HXR Imaging spectroscopy constraints

CdTe or CdZnTe?

Fine pitch imaging spectrometers : two recent examples for fair comparisons ...

**Outro : Take home message** 





This presentation is not a review of all CdTe and CdZnTe imaging spectrometer technologies worldwide. It shows some advanced developments illustrating different variants in shape, technolgogies, usage and performance of both materials with pixelated patterns, mostly in the field of space science.

A couple of examples are emphasized because they have been using both materials with the exact same electronics, allowing kind of a fair comparison between those two materials when they are used the same way.

The presentation is not intended to review all technologies and inevitably misses some important developments, including in space science.

# Cea IMAGING SPECTROSCOPY



#### Coded Mask:

- Wide FoV
- Modest Angular resolution
- High Dynamic range in E

### Grazing incidence mirrors

- High Angular resolution  $\bullet$
- Narrow FoV  $\bullet$
- Limited energy range  $\bullet$











## **CONTINUES A BETTER SENSITIVITY – DIRECT IMAGING**



NuStar / NASA



Huge constraints on the detector design

 $10^{-5}$ 

10<sup>-6</sup>

10<sup>-7</sup>

10<sup>-8</sup>

10<sup>-9</sup>

0

60

Exposure 1 Ms -  $3\sigma$  detection -  $\Delta E = E/2$ 

1 ACrob

20

40

Energy (keV)

Integral ISGRI

1 mCrab

Astro-H

NuSTAR

COSPIX

80

100

### WHY SHOULD WE USE Cd(Zn)Te COMPOUNDS?

- High Z
- High density
- High resistivity
- Moderate cooling
- Compact design
- Flexible pixelization
- Space proven





# Cea A WORLDWIDE CHALLENGE – FEW RACERS

# Fine pitch



CEA/CALISTE - MACSI 670 eV fwhm @ 60 keV 32x64 pixels 625 µm

### Limousin+14



4x(32x32) pixels 605 μm

### Harisson+13



# Ce A WORLDWIDE CHALLENGE – FEW RACERS

# Very Fine pitch



RAL/HEXITEC ~1 keV fwhm @ 60 keV 4x(80x80) pixels 250 µm

Jowitt+21



JAXA/HXI ~2 keV fwhm @ 60 keV 128x128 pixels 250 µm

Hagino+21



CEA/MC2 ~ 0.95 keV fwhm @ 60 keV 32x32 pixels 250 µm

### Allaire+23



# A WORLDWIDE CHALLENGE – FEW RACERS

# Ultra Fine pitch



### Smolyanskiy+24



#### Furukawa+19

# **KEY PARAMETERS FOR SPECTROSCOPY**

# Signal

### Collect efficiently the charge carriers

- + HV
- + Thin crystals
- + Corrections and 3D

# Noise

- In CdTe Fano (0.15) limit is
   + 147 eV FWHM at 5.9 keV
   + 226 eV FWHM at 13.9 keV
   + 467 eV FWHM at 59.54 keV
- Minimize Dark currentStray capacitance





### SMALL PIXEL EFFECT

# **Signal Induction**



$$CCE = \frac{Q}{q_0} = N \cdot \left[ \left( \frac{d}{\mu \tau \cdot V} \right) \cdot \int_0^d \varphi_0(x) \cdot e^{-\frac{d \cdot x}{\mu \tau \cdot V}} \cdot dx + e^{-\frac{d^2}{\mu \tau \cdot V}} \right]$$

# **Nearly Single carrier detectors**



# CdTe Vs. CdZnTe ?

CdTe

- Band gap 1.44 eV
- Ionization energy ~4.42
- Z = 50
- D = 5.85 g/cm<sup>3</sup>
- ~1E9 Ω.*cm*
- "Easy" to procure in Space grade
- $\mu_{\rm e} \tau_{\rm e} \sim 3 \times 10^{-3} \, {\rm cm}^2 {\rm V}^{-1}$  (Acrorad)
- $\mu_h \tau_h \sim 2 \times 10^{-4} \text{ cm}^2 \text{V}^{-1}$  (Acrorad)

# CdZnTe

- Band gap ~1.62 eV
- Ionization energy ~ 4.6 eV
- Z = 50
- D = 5.8 g/cm<sup>3</sup>
- 1E10 ~1E11 Ω. cm
- $\mu_{e} \tau_{e} \sim 13 \times 10^{-4} \text{ cm}^{2} \text{V}^{-1} \text{ (Kromek)}$
- $\mu_e \tau_e \sim 100 \times 10^{-3} \text{ cm}^2 \text{V}^{-1} \text{ (RedLen)}$
- $\mu_{e} \tau_{e} \sim 11 \times 10^{-3} \text{ cm}^{2} \text{V}^{-1} \text{ (RedLen HF)}$
- $\mu_h \tau_h \sim 5 \sim 8 \times 10^{-5} \text{ cm}^2 \text{V}^{-1}$



### **Cerror Other consideration for fair comparisons ....**





# Both technologies evaluated with CdTe and CZT



RAL/HEXITEC ~1 keV fwhm @ 60 keV 4x(80x80) pixels 250 µm

Lowitt+21



CEA/CALISTE - MACSI 670 eV fwhm @ 60 keV 32x64 pixels 625 µm

Limousin+14

### Caliste-256 technology



Am 241 spectrum obtained with Caliste 256 SN5 equipped with a CdTe Schottky diode summing the 256 individually calibrated spectra (0 °C, 300 V, 9.6  $\mu$ s peaking time). Energy resolution at 59.54 keV is 0.85 keV FWHM. Zoom in the 10–25 keV band emphasizes the energy resolution revealing the structure of the 17.75 keV triplets.



Am 241 spectrum obtained with Caliste 256 SN4 equipped with a CZT detector summing the 256 individually calibrated spectra (-15 °C, 800 V, 6.0 µs peaking time). Energy resolution at 59.54 keV is 1.09 keV FWHM. Zoom in the 10–25 keV band emphasizes the energy resolution revealing the triangle shape of the 17.75 keV triplets.

#### Limousin+11

### C22 Caliste Technology best spectra with CdTe



Cez

# HEXITEC Technology



UK Research and Innovation



# Single events only Lowitt+21

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DE LA RECHERCHE À L'INDUSTRIE HEXITEC Technology **UK Research** Science & Technology and Innovation **Facilities Council** CdTe ---- Isolated — Adjacent Bipixels — Diagonal Bipixels HF – CdZnTe 0.5 Frame occupancy: 0.4% Frame occupancy: 0.4% (a) CdTe Frame occupancy: 0.4% 0.92 keV FWHM 0.4 60 0.3 4.85 keV FWHM 0.2 50 (PDF) 0.1 **Adjacent Bipixels** 0.0 (b) HF – CdZnTe Counts 0.95 keV FWHM 40 Frame occupancy: 0.4% 0.4 1.63 keV FWHM -> 0.3 - 1.68 keV FWHM Normalised (c) 30 - **(a)** 0.2 0.1 0. (c) HF – CdZnTe — γ: 59.5 keV Frame occupancy: 4% 20 0.4 0.3  $\gamma$  pileup peak: 0.2 119 keV Fluorescence pileup peaks: 10 82 - 96 keV 0.1

Noise threshold: 3 keV

0.0

50

60

70

80

Photon energy (keV)

90

100

120

130

0

110

# Multiple events Koch-Mehrin+21

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### OUTRO: take home message

And the winner is ...
CdTe is cheaper than CZT
CdTe is easy to procure
CdTe show very good performance at low temperature
CdTe show very stable performances in space
CdTe Schottky dark current is unbeatable
Ideally, CdTe would win with more stability in time ...

### Or ...

CZT is more stable in time than CdTe Schottky at low flux

- CZT is good at room temperature
- CZT is ideal when more than 2 mm thick detectors are needed
- CZT HF is looking good in inter pixel gaps

# Cea THANKS FOR YOUR ATTENTION !



CEA/MC2 - MACSI 4x(32x32) pixels 250 µm

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