

SFB 1258

Neutrinos
Dark Matter
Messengers



BAW

BAYERISCHE
AKADEMIE
DER
WISSENSCHAFTEN

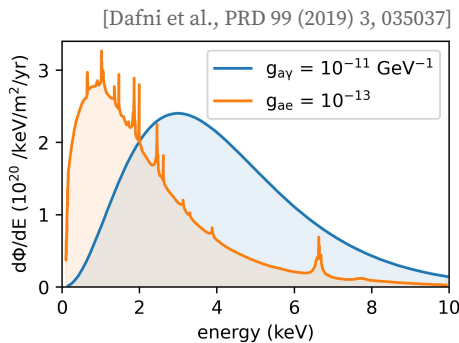
Towards low-background silicon drift detectors for

Joanna Bilicki, Patrick Bongratz, Frank Edzards, Susanne Mertens, Lucinda Schönfeld, Juan Pablo Ulloa Beteta, Korbinian Urban, Christoph Wiesinger, Michael Willers () , XDEP conference, 05.02.2024

UXO detector requirements

- search for **solar-axion induced X-rays**

- 1) sufficient **energy threshold** and **resolution**

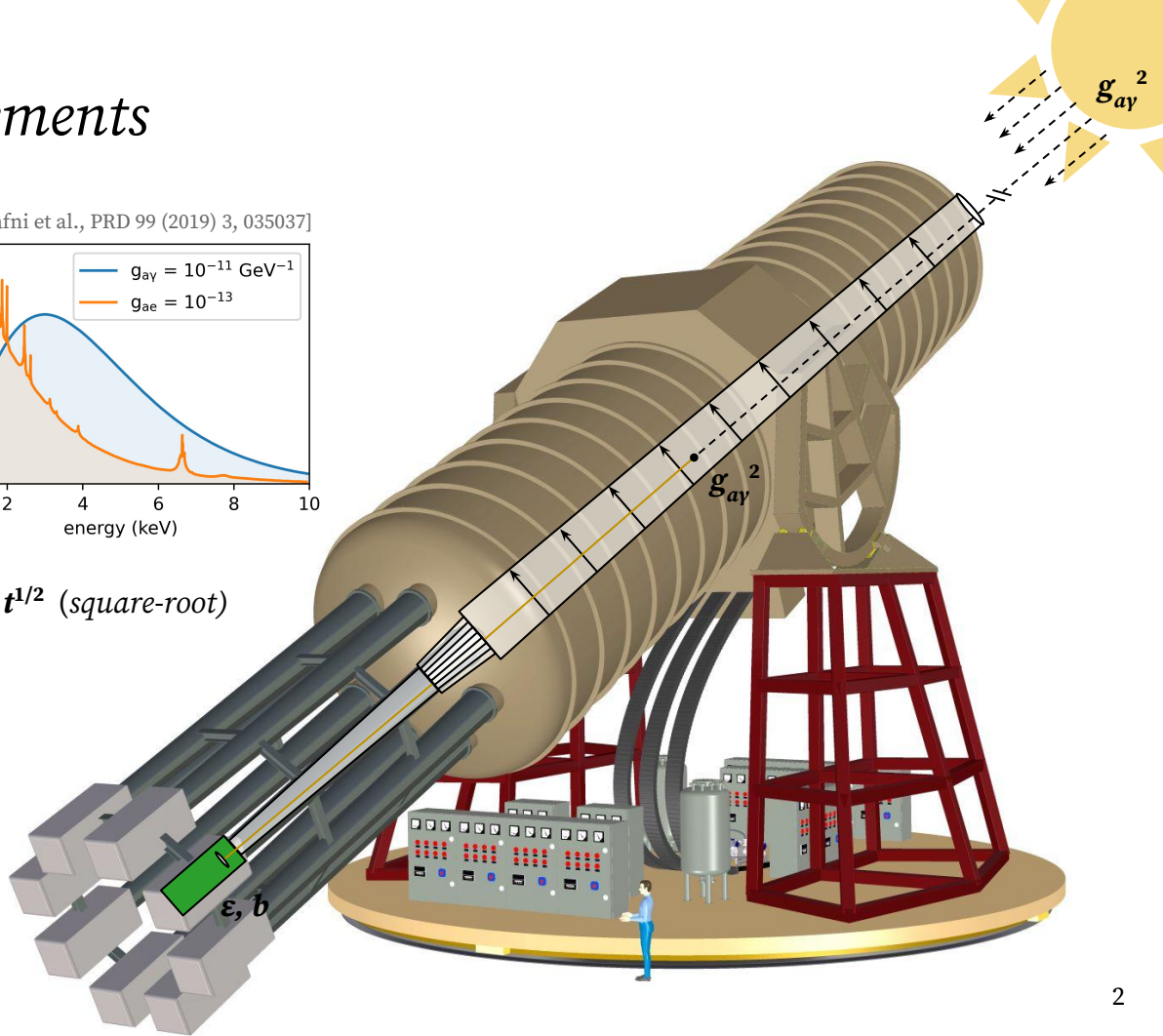


→ **sensitivity** scales with $g_{ay}^{-4} \propto \epsilon / b^{1/2} \cdot t^{1/2}$ (square-root)

- 2) high X-ray **detection efficiency** (ϵ)

- 3) **ultra-low background** (b)

- goal **10^{-8} cts/keV/cm²/s**
(single events per year)



Background challenge

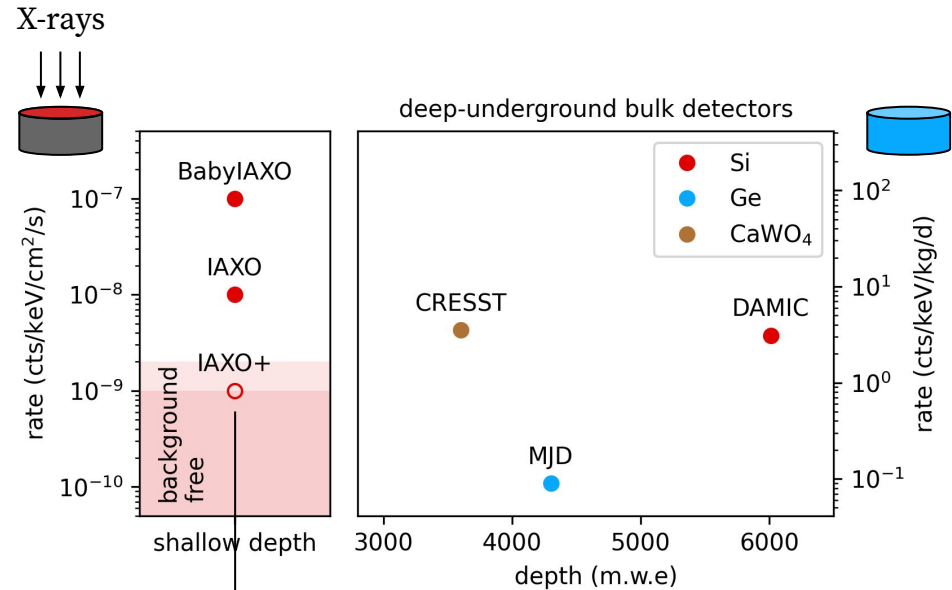
- **similar to** other rare-event searches, e.g. **CEvNS**, **WIMP**, **$0\nu\beta\beta$ decay** searches

→ radiopurity, **background mitigation**

but

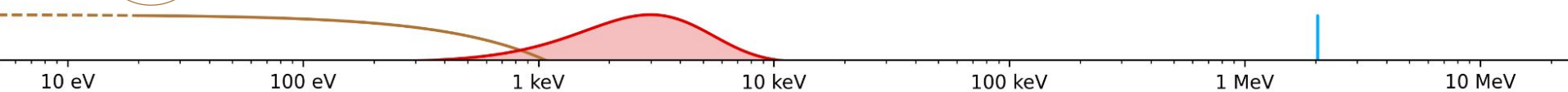
- deep-underground like performance **at shallow depth**

→ **cosmic-ray** induced backgrounds



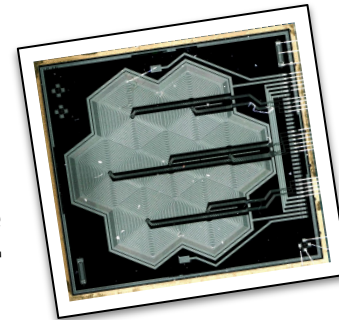
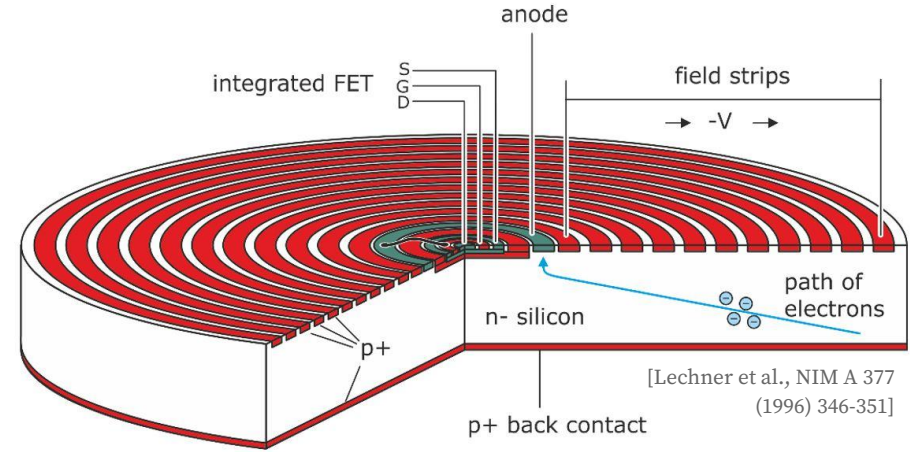
$$g_{ay}^{-4} \propto \epsilon_d \cdot t$$

(linear)



Silicon drift detectors (SDDs)

- tiny read-out electrode, **low capacitance**
- low **energy threshold** (< 1 keV) ✓
and good **resolution** (< 200 eV FWHM at 6 keV) ✓
- thin deadlayer (< 100 nm), **no entrance window**
- high X-ray **detection efficiency** in [1, 10] keV ✓
- **semiconductor-grade materials**, little auxiliaries
- great potential for **low-background** operations □



7x3-mm pixel **TRISTAN** prototype
SDD array with integrated JFET

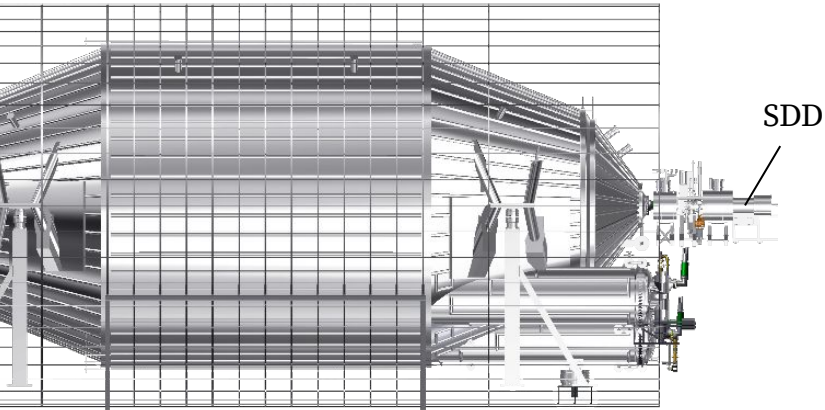
TRISTAN



*poster by
Christian
Forstner*

keV-sterile neutrino search with KATRIN

[Mertens et al., J.Phys.G 46 (2019) 6, 065203]



- **high-rate electron spectroscopy**
- ultra-high vacuum compliance, calibration

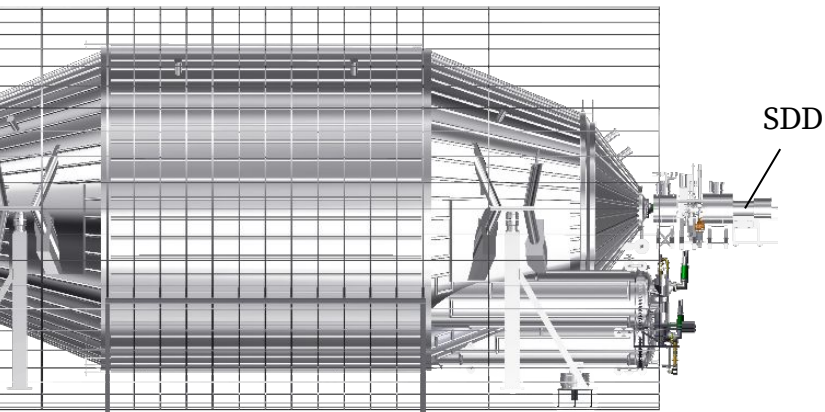
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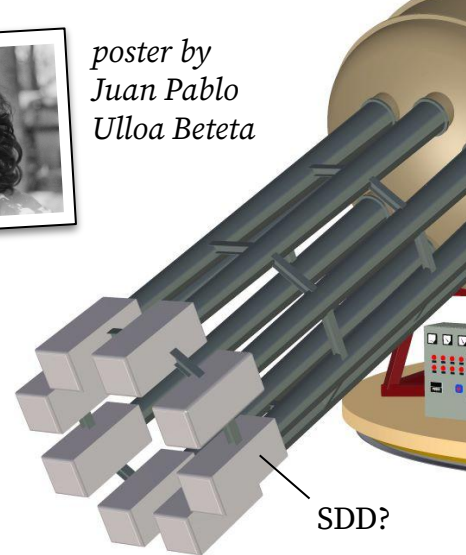


poster by
Juan Pablo
Ulloa Beteta

solar axion search

[Armengaud et al., JINST 9 (2014) T05002]

- **rare-event X-ray detection**
- ultra-low background,
high-efficiency



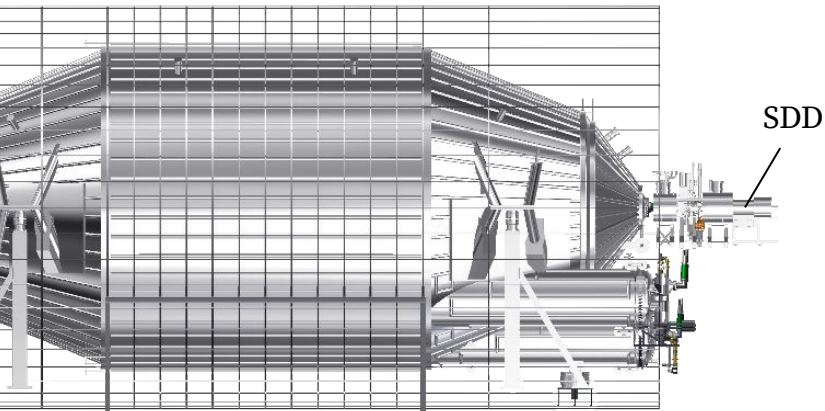
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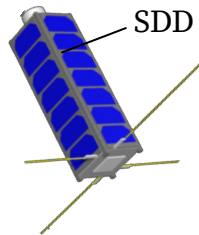
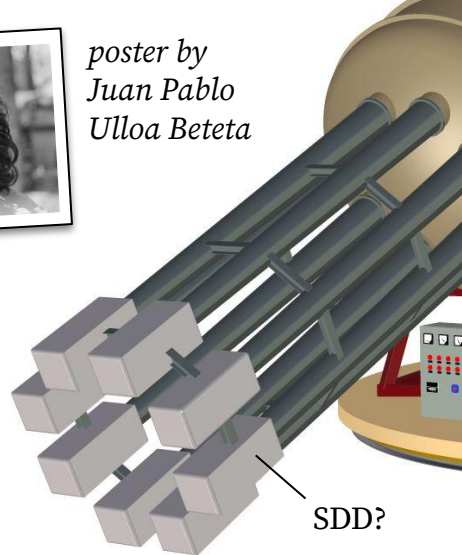


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ComPol

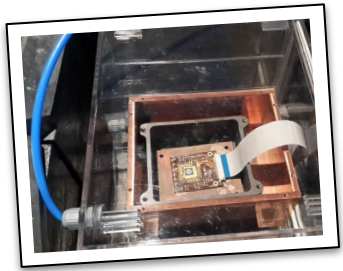
X-ray polarization measurement of Cygnus X-1

- compact **Compton telescope**, CubeSat mission
- space environment, radiation hardness, remote operation

poster by
Jonas Schlegel



TAXO development steps



SDD pathfinder

material insufficiencies,
close-by electronics

→ $(1.9 \pm 0.2) \cdot 10^{-5}$
cts/keV/cm²/s
at shallow depth

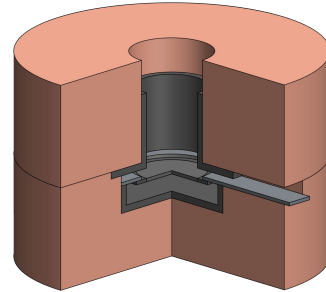
[Houdy et al., 2019]



TAXO demonstrator

consequent **low-background**
approach, **far electronics**

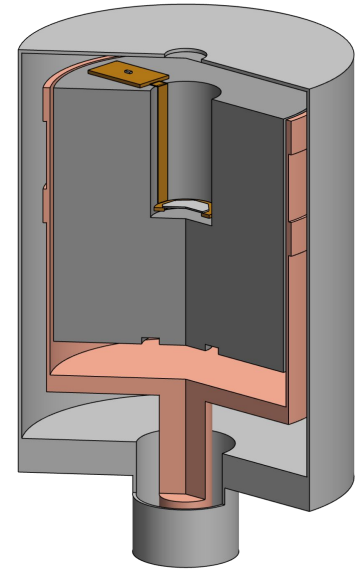
→ demonstrate
 $O(10^{-7})$ cts/keV/cm²/s



BabyTAXO

dedicated SDD production,
improved **passive shield**

→ target
 $< 10^{-7}$ cts/keV/cm²/s
at shallow depth



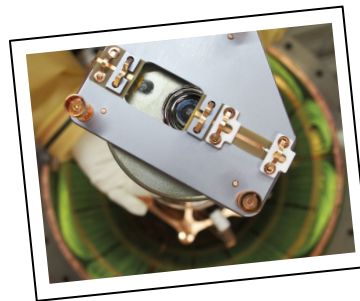
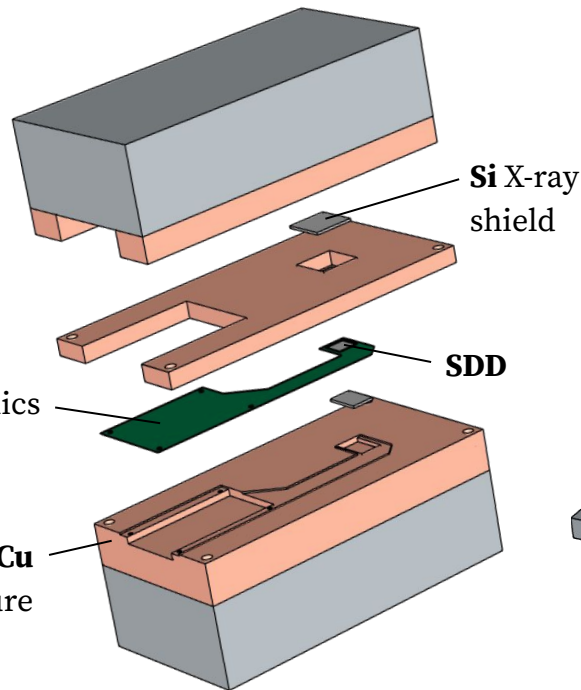
TAXO

all-semiconductor **active shield**
SDD-in-HPGe detector

→ target
 $< 10^{-8}$ cts/keV/cm²/s
at shallow depth

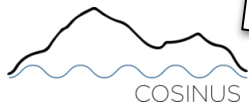
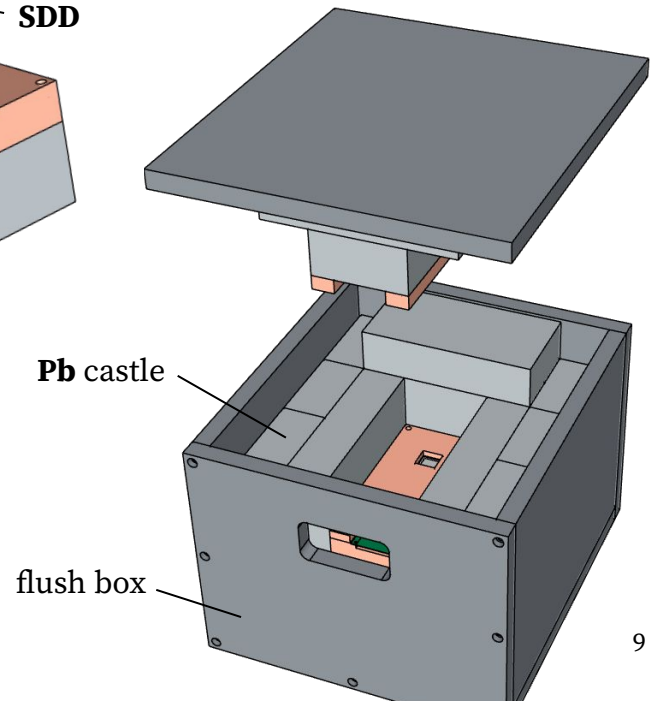
TAXO demonstrator

- **passive shield** setup with
 - prototype **SDD** on polyimide board, far electronics
 - **Si** X-ray shield, radiopure **Cu** enclosure
 - compact **Pb** castle, flush box



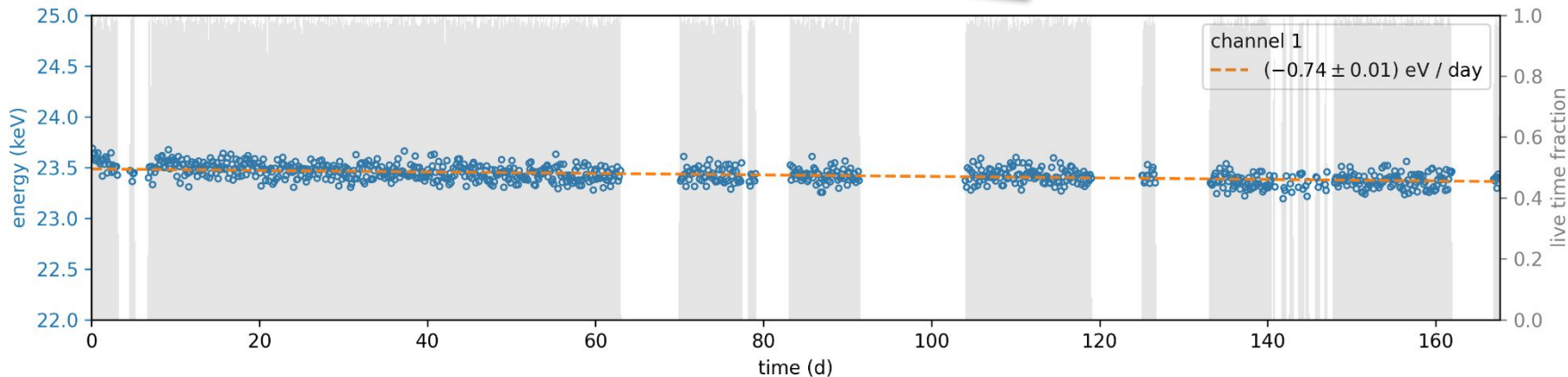
→ demonstrate $O(10^{-7})$ cts/keV/cm²/s

- 1) **deep underground** at Canfranc
- 2) at **shallow depth** with muon veto and neutron shield



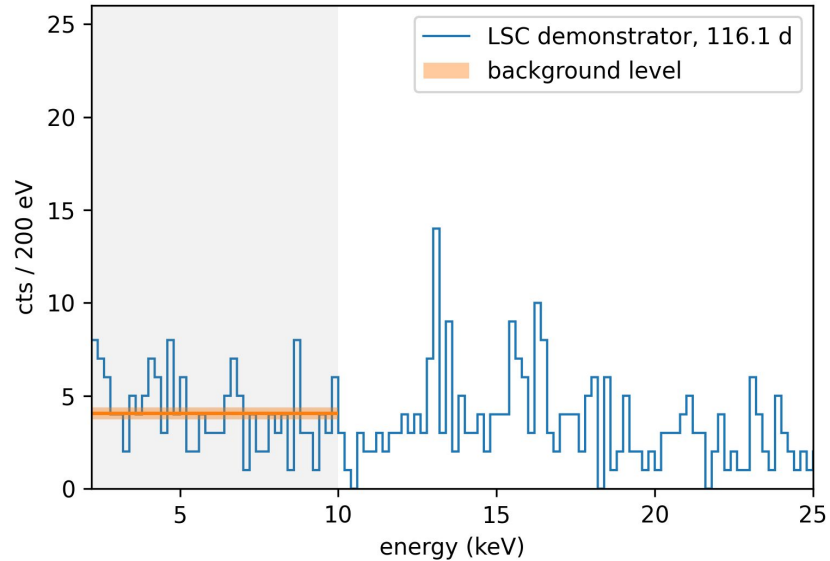
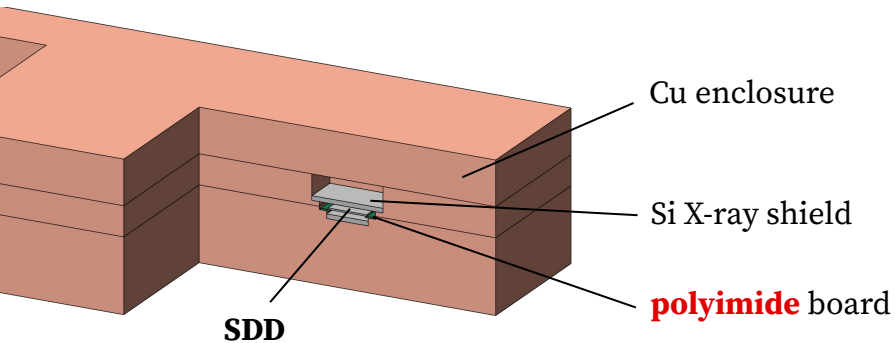
TAXO demonstrator at Canfranc

- installed in Lab2500 (2450 m.w.e.)
- determination of intrinsic **non-cosmic background**
- stable **long-term operation**, **gain monitoring**
- **> 100 d** live time, **< 1 eV/d** gain drift



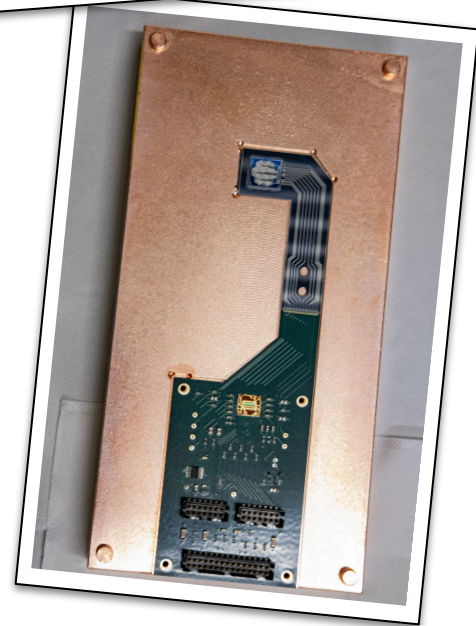
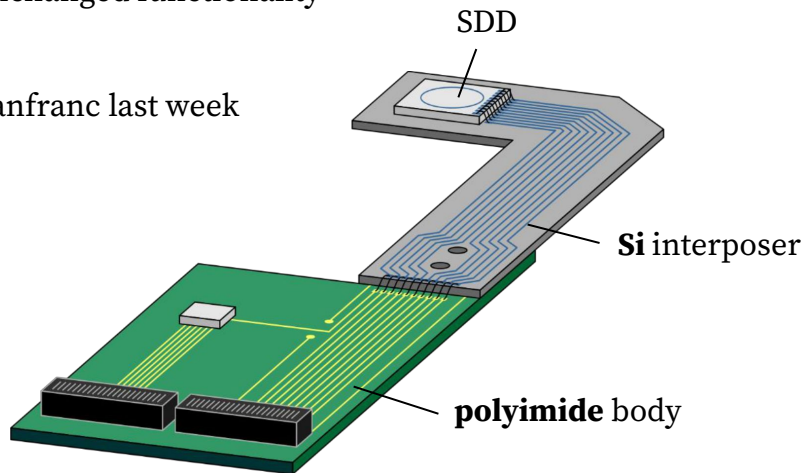
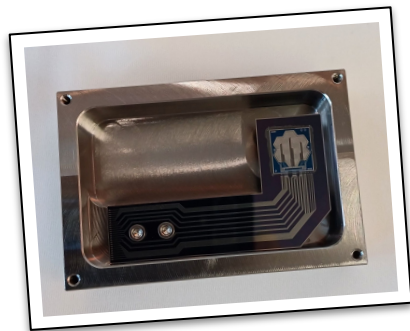
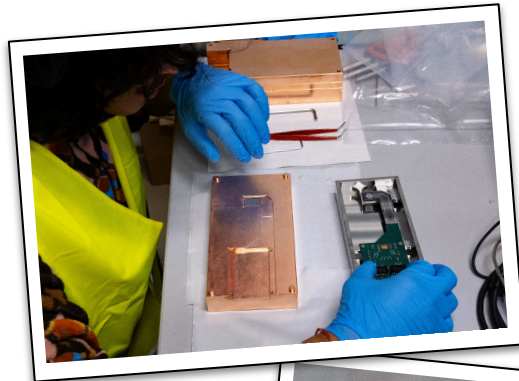
TAXO demonstrator at Canfranc

- noise situation suboptimal
- **2 keV** (online) **threshold**, 300 eV FWHM at 10 keV
- **background level higher than anticipated**
- $(5.6 \pm 0.5) \cdot 10^{-6}$ cts/keV/cm²/s in [2, 10] keV
- **X-ray lines** from close-by contaminants
- **polyimide** assay, **activity 10-100x above expectations**



TAXO demonstrator upgrade

- polyimide board **head area dominates** background budget
- split detector board into
 - **Si interposer head**, improved radiopurity in SDD vicinity
 - **polyimide body**, unchanged functionality
- first upgrade attempt at Canfranc last week



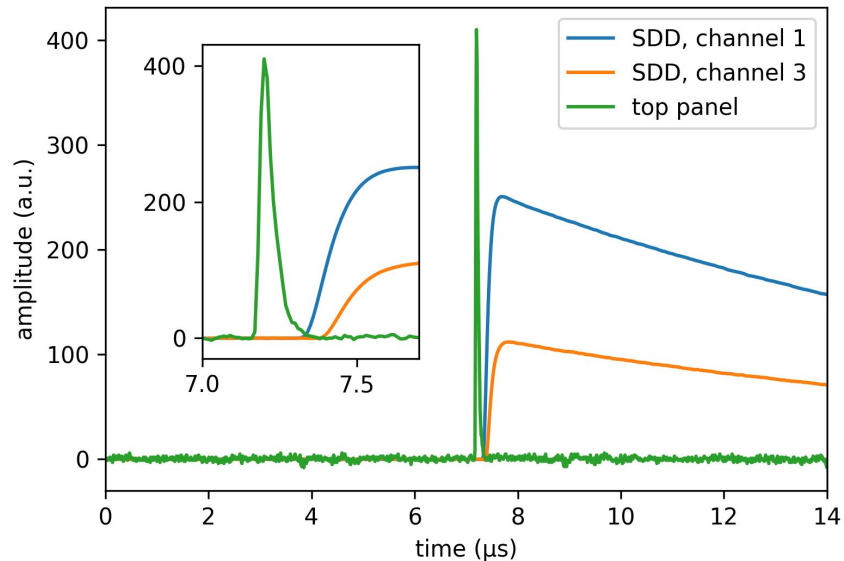
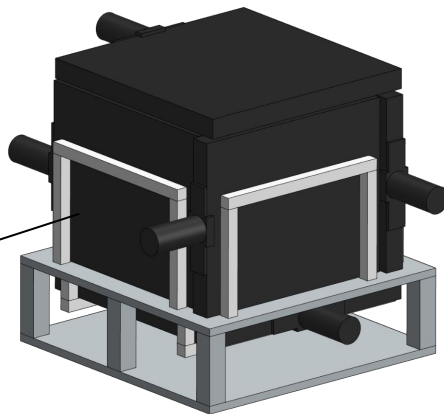
TAXO demonstrator at TUM


- **twin** of Canfranc demonstrator with
 - **muon veto**, 6 plastic scintillator panels
 - **neutron shield**, borated polyethylene
- disentangle **cosmic-ray induced background**

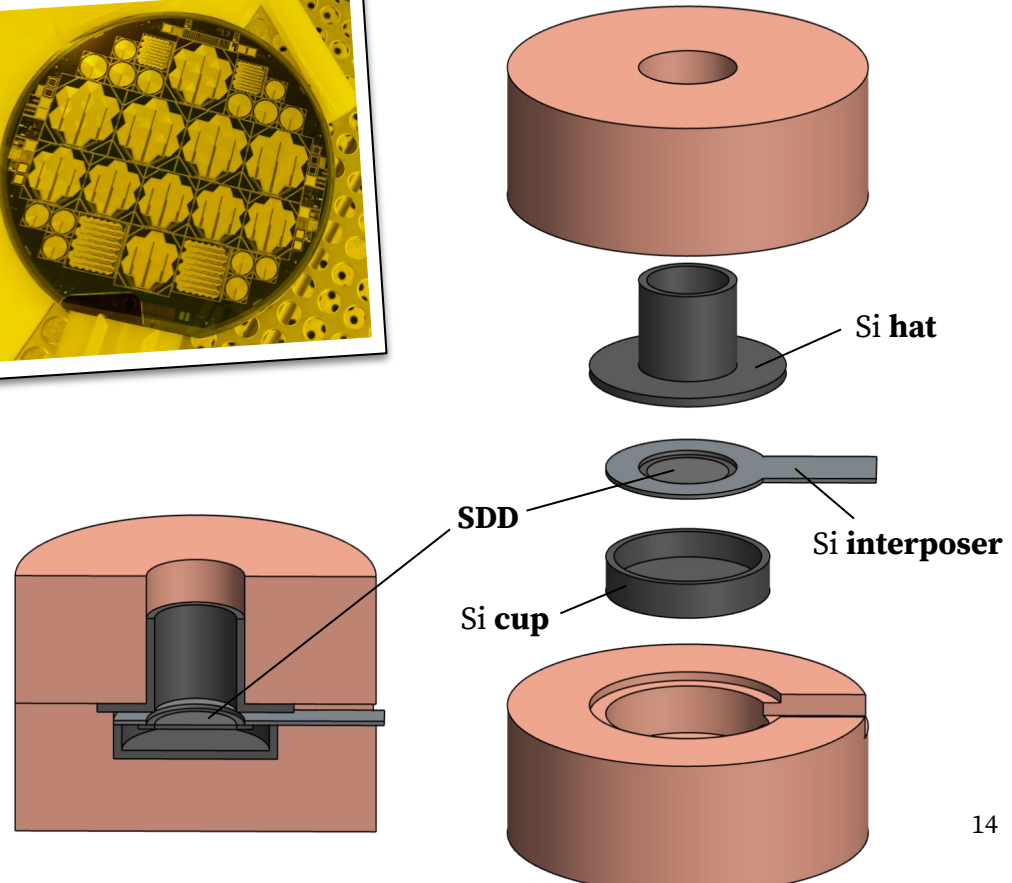
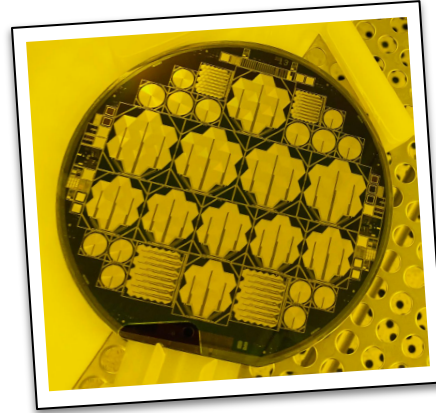
- background measurements in preparation
 - **above ground** at TUM
 - **at shallow-depth** at DESY
HERA south hall



scintillator panel

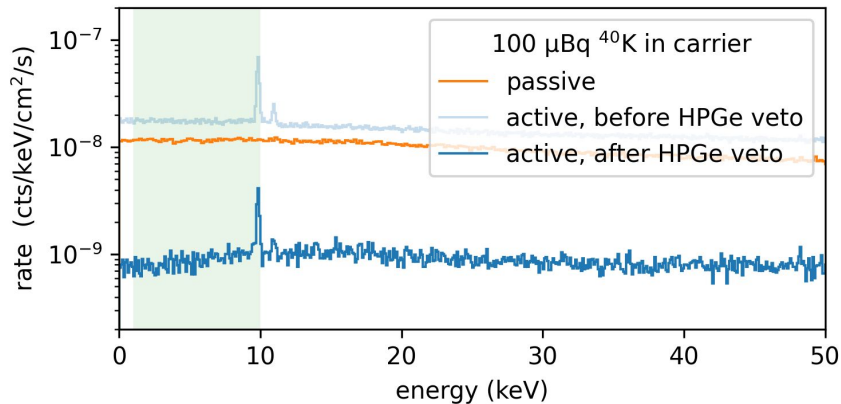


- **passive shield** setup, based on demonstrator experience
 - maximal **X-ray shield coverage**
 - radiopure **interposer**
 - optimized **neutron shield**
 - dedicated **large-area SDDs**
 - beam tube, **vacuum** interface
- target $< 10^{-7}$ cts/keV/cm²/s, potential detector for **Baby  X₀**



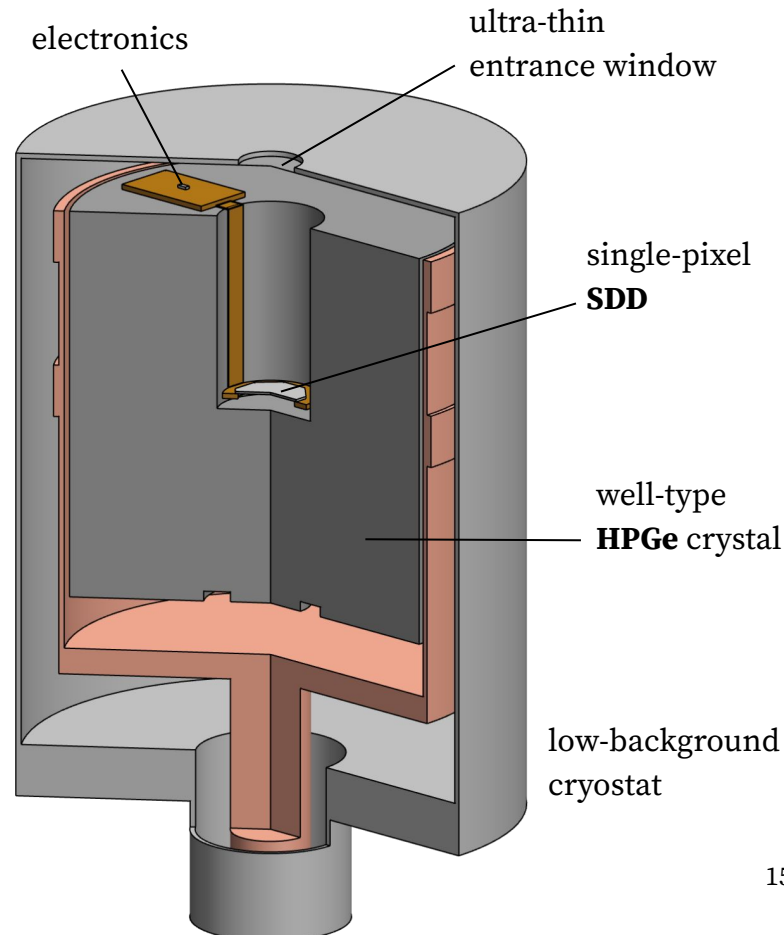
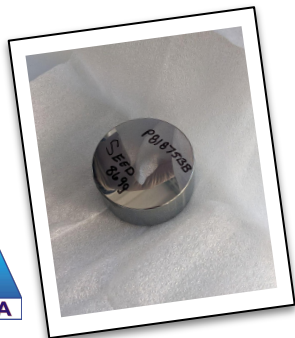
TAXO - SDD-in-HPGe detector

- all-semiconductor **active-shield design**, ultra-pure $\sim 4\pi$ veto



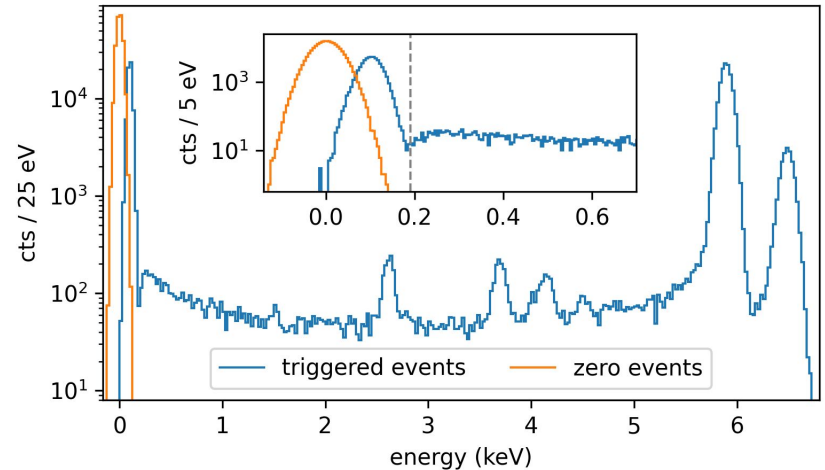
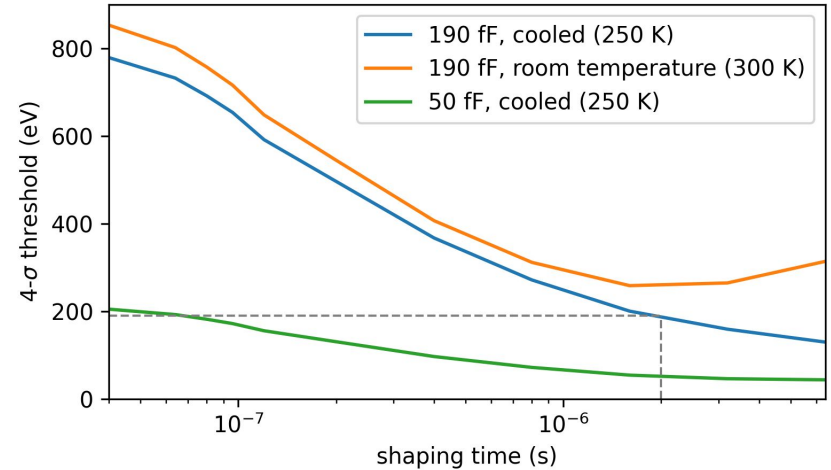
- well-type HPGe detector **prototype**, **proof-of-principle** planned for this year

→ target $< 10^{-8}$ cts/keV/cm²/s,
potential detector for TAXO

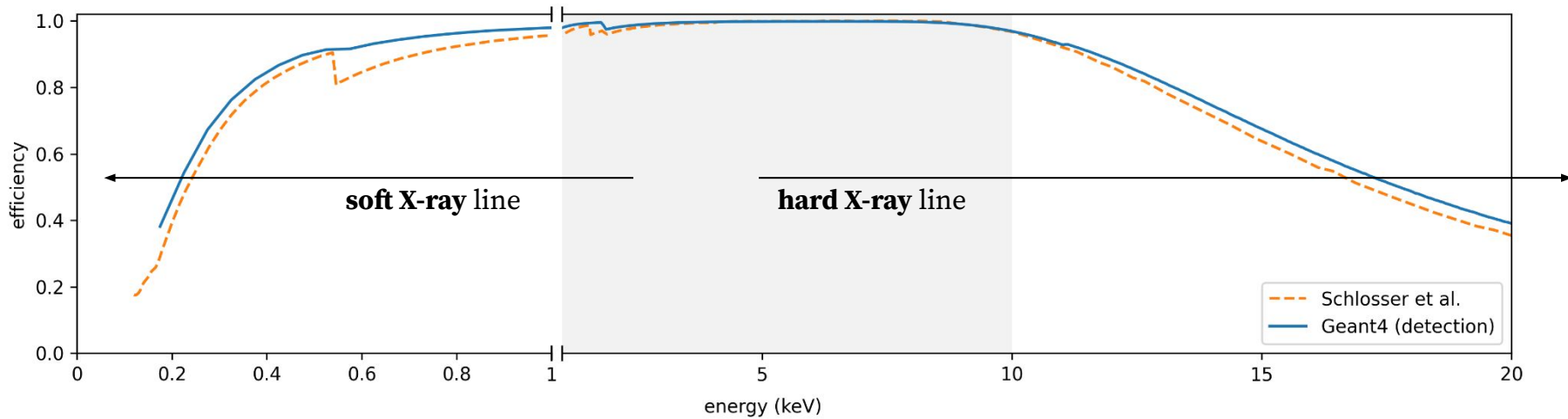



Energy threshold

- depends on **noise performance**
 - read-out capacitance
 - temperature
 - trigger filter (shaping time)
- **200 eV** achieved with TRISTAN SDD (cooled, 2 μ s shaping time)
- **25 eV** reported in literature, droplet SDD [Strüder et al., Microscopy Today 28 (2020)]
- **sub-keV threshold** feasible, but cooling/close-by electronics could alter background performance





Detection efficiency



- **450- μm SDD with 50 nm dead layer**, no entrance window
 - **> 95%** in [1, 10] keV, but also substantial for exotic axion models, **> 50%** above **250 eV**, **> 65%** at **14.4 keV**
- efficiency measurement at  **metrology beamline** in preparation

Conclusions

- **TAXO** SDD development to meet challenging  X-ray detector requirements with a
 - conventional **passive shield** setup
 - novel all-semiconductor **SDD-in-HPGe active-shield** design
- **background exploration** deep-underground at Canfranc and at shallow depth **ongoing**
- efficiency measurement at  **metrology beamline** in preparation
- application in **other particle physics projects**
 - keV-sterile neutrino search with **TRISTAN** in KATRIN
 - X-ray polarization measurement with **ComPol**

Backup

Background projections

- stand-alone **Geant4** application, demonstrator and different shielding geometries implemented
- **radiogenic** backgrounds
 - **assay** results, radiopurity **estimates**
 - **cosmogenic activation** (^3H , ^{32}Si)
[Saldanha et al., PRD (2020) 10, 102006;
Orrell et al., Astropart.Phys. 99 (2018) 9-20]
- **cosmic** backgrounds
 - full (correlated) above ground **cosmic ray flux**, cosmic ray library (**CRY**)
[https://nuclear.llnl.gov/simulation/doc_cry_v1.7/cry.pdf]
 - significant **reduction of neutron-induced background** with borated polyethylene

